

The Hunt for Dark Matter

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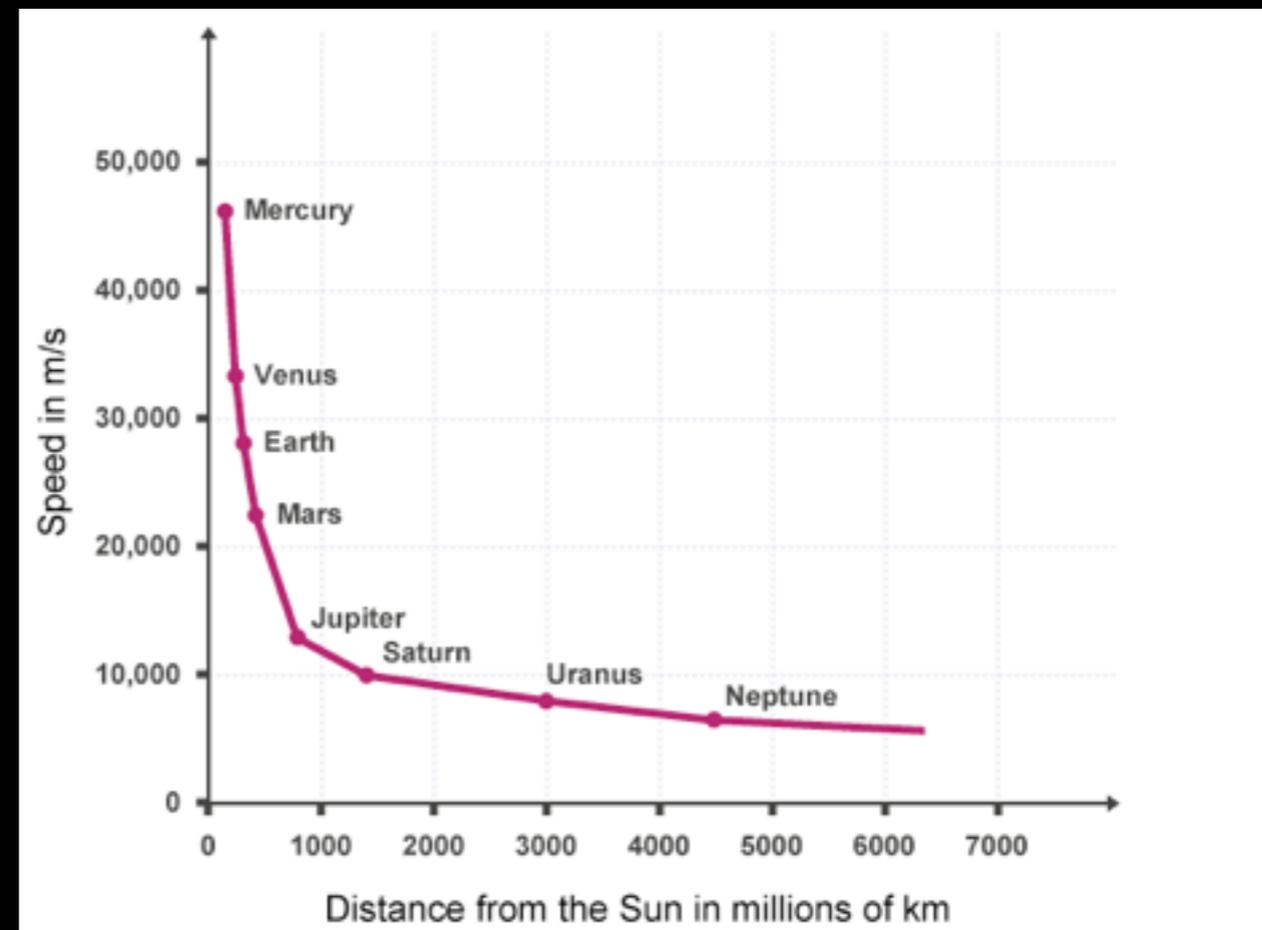
What is Dark Matter?

- **Matter is anything that has mass**
- **All matter attracts through gravity**



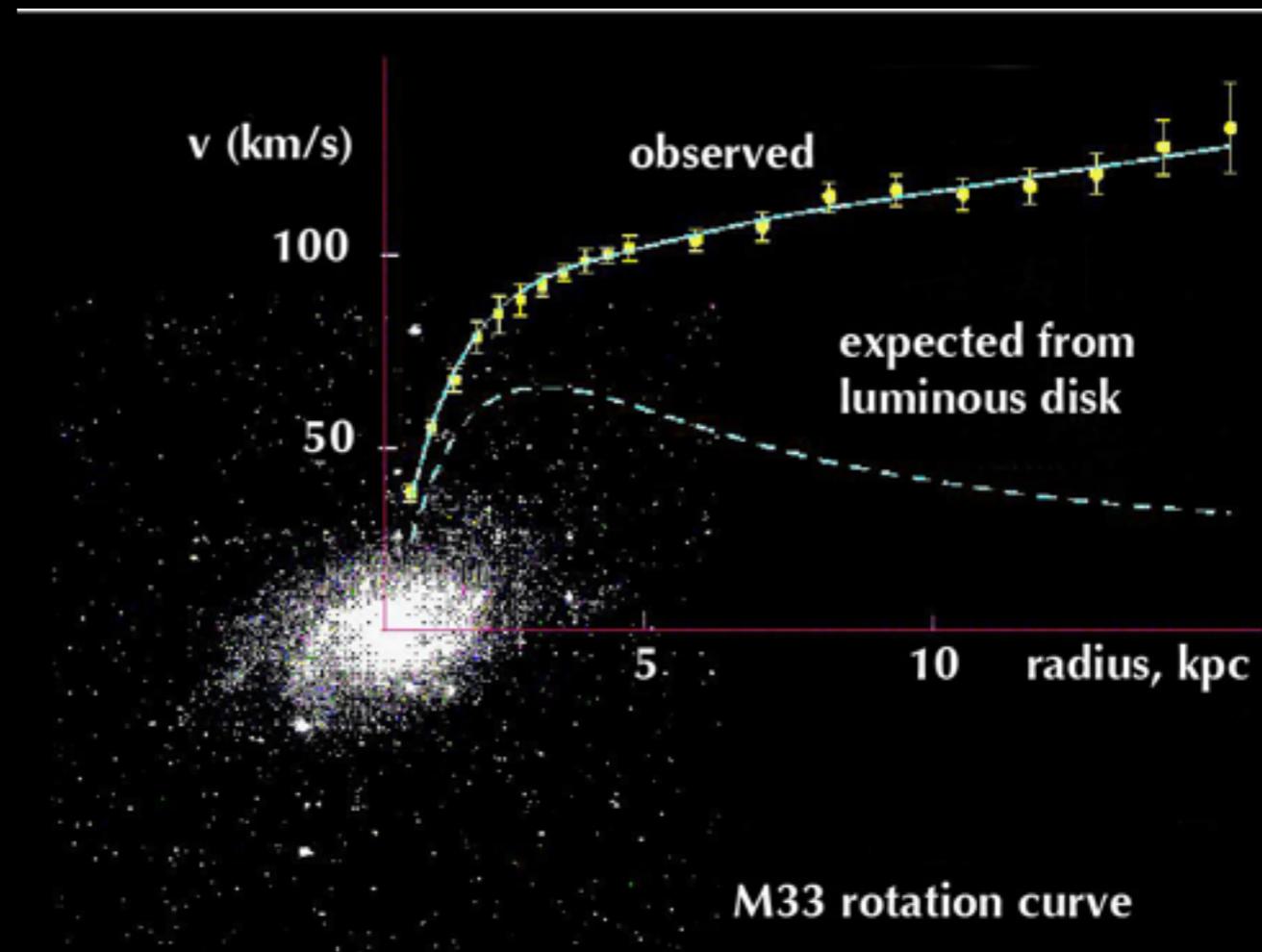
What is Dark Matter?

- Newton showed that the strength of the gravitational force decreases with increasing distance
- This model works very well for our solar system



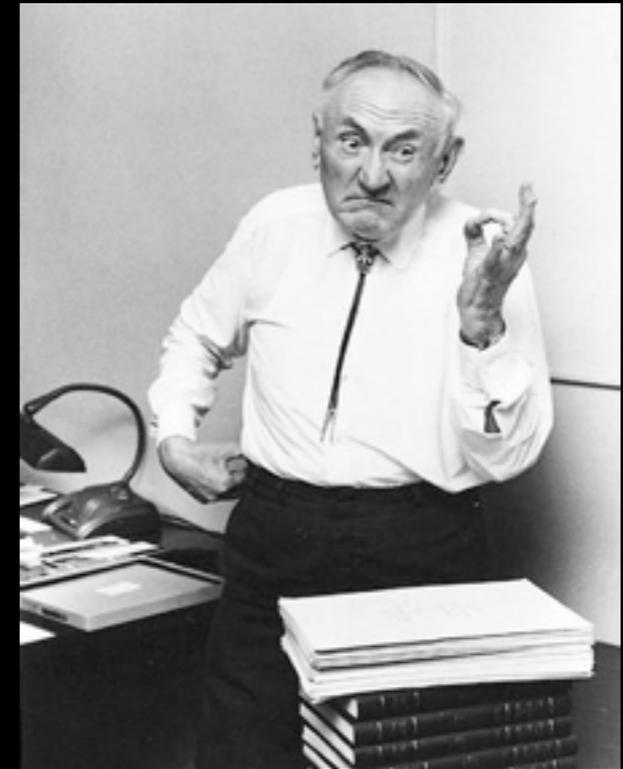
What is Dark Matter?

- However, when we apply this to galaxies, the model fails.
- This is especially true at distances very far from the center of the galaxy.



What is Dark Matter?

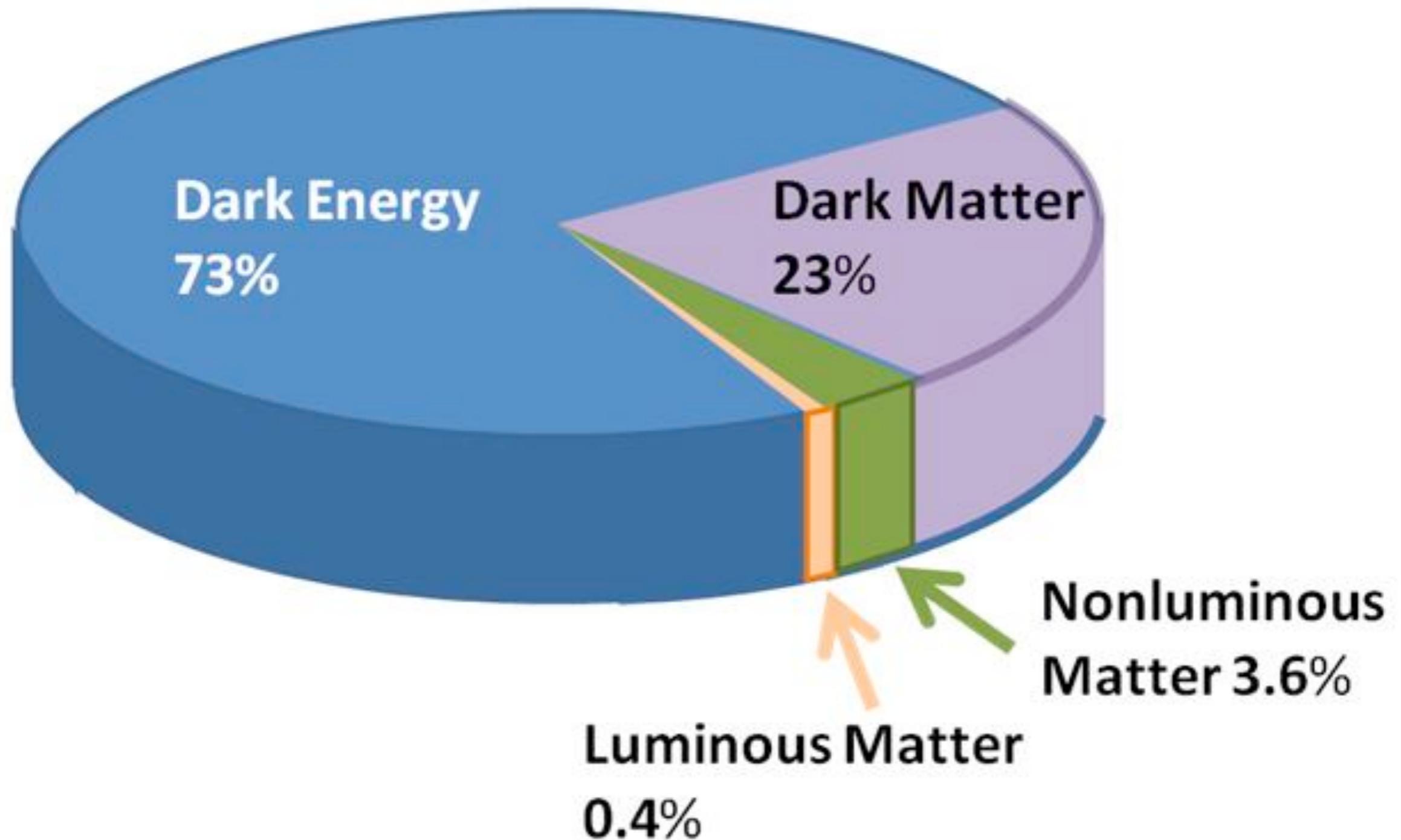
- This anomaly was first pointed out by Fritz Zwicky . He investigated clusters of galaxies revolving around each other.
- Vera Rubin greatly improved the quality of the data in the 1970s, convincing mainstream astronomers that the effect was real.



What is Dark Matter?

- **We believe that Newton's Laws apply, even to large systems**
- **Thus we believe that there is a new mass, which we are not including in our calculations.**
- **We can't see the matter that is creating this mass.**
- **Because we are creative we call this matter "dark matter"**

How Much Dark Matter?



Detecting Matter

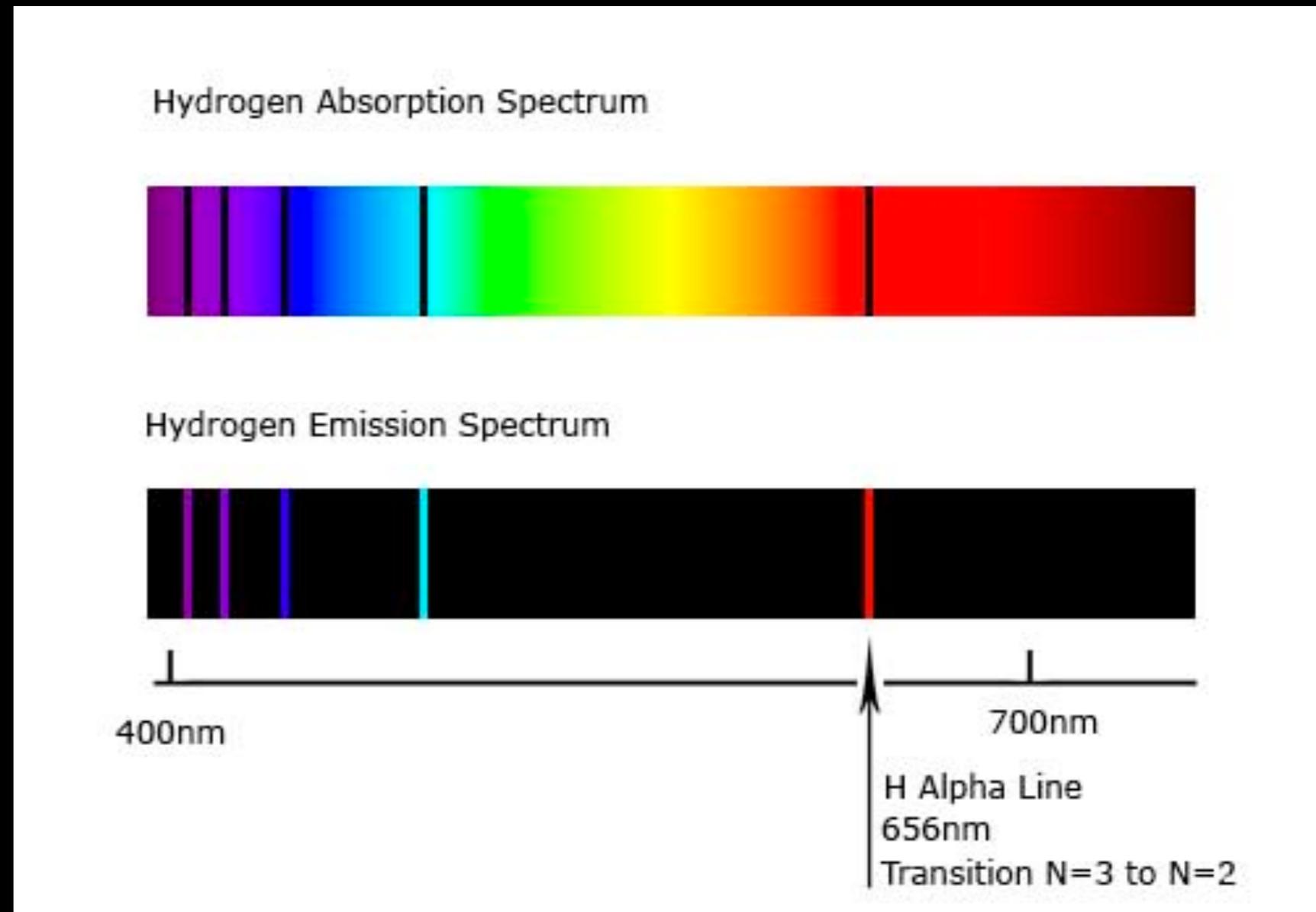


Light Emitters



Light Absorbers

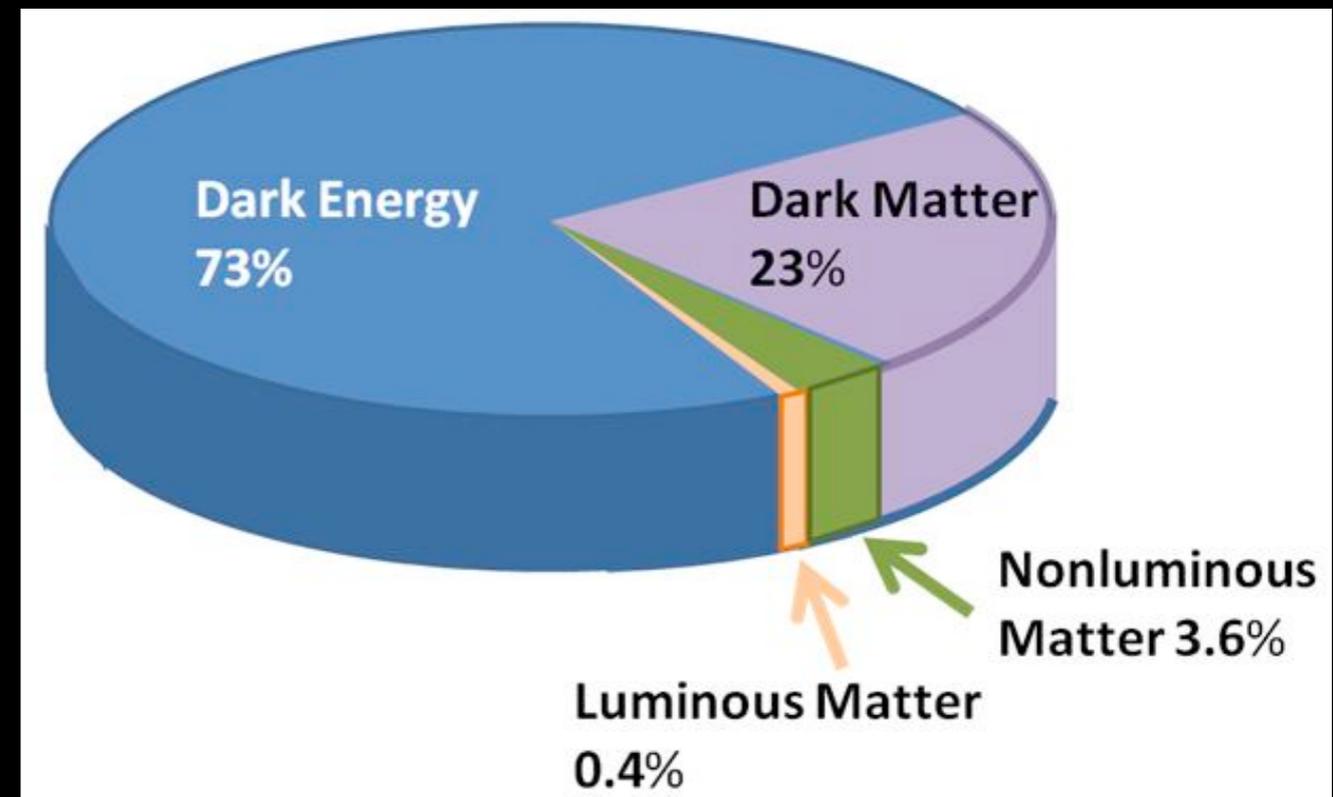
Detecting Light Absorbers



Matter absorbs light more effectively at specific frequencies

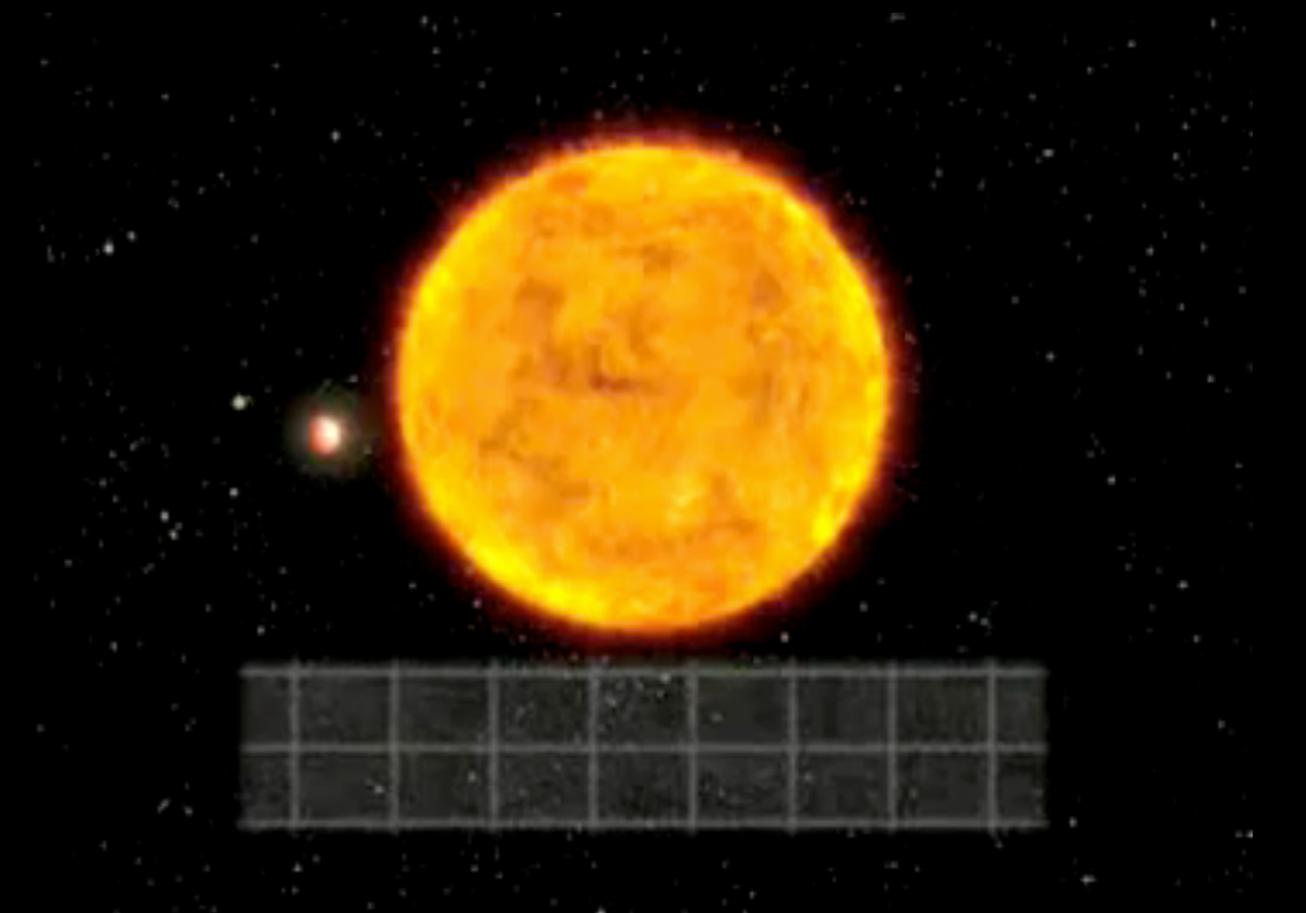
Planets

- Planets are massive and don't emit much light.
- However - there is 50x as much dark matter as stellar matter
- Jupiter is 1/10,000x as massive as the Sun.
- We need 500,000 Jupiter mass planets for each star in the galaxy.



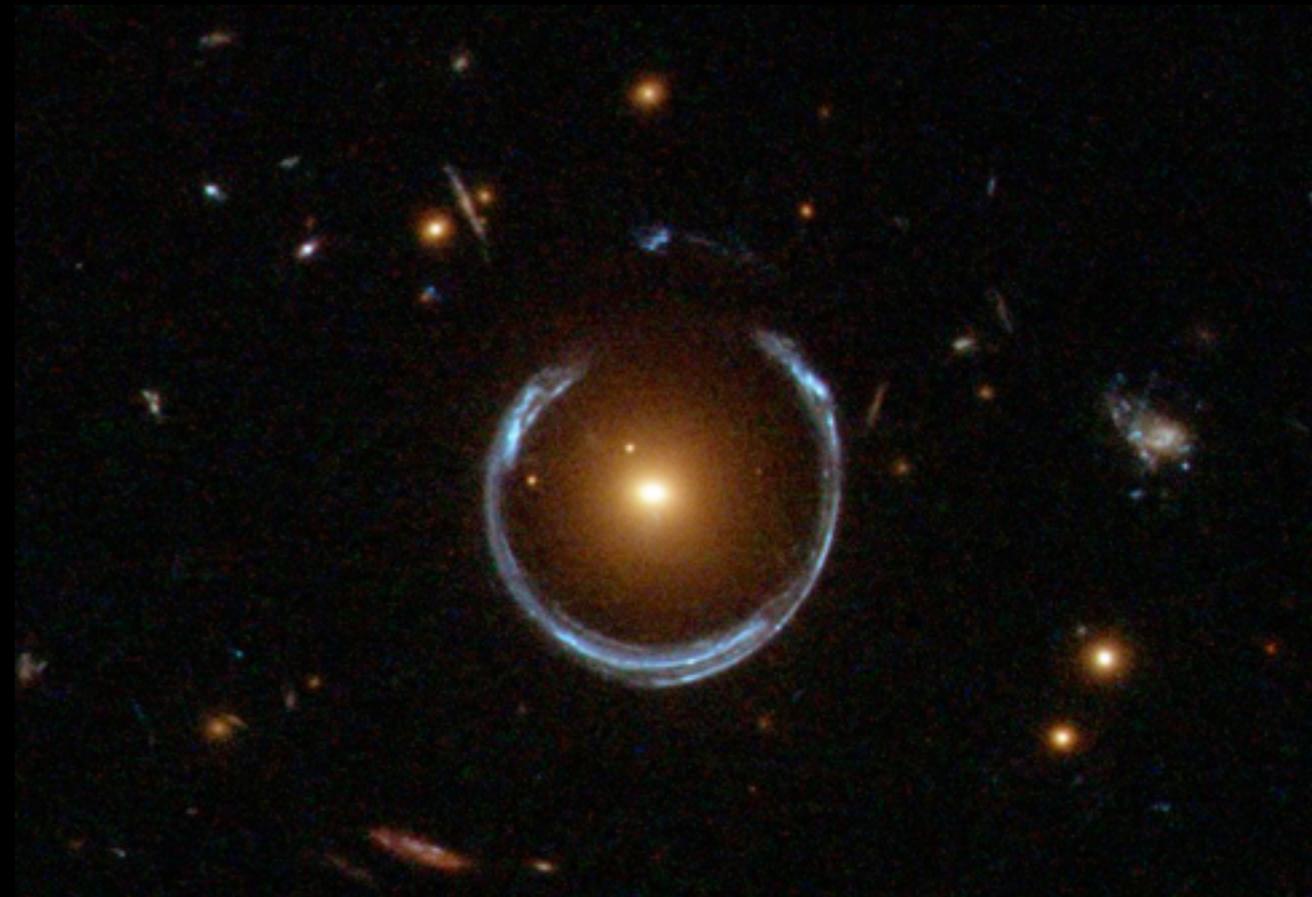
Planets

- **We can detect these planets if they happen to move in front of stars and block the starlight**



Black Holes

- **Black holes can be much more massive, we don't need as many**
- **They are small, so they don't block much light.**
- **But black holes can warp the light behind them, creating detectable patterns in the sky.**



Black Holes



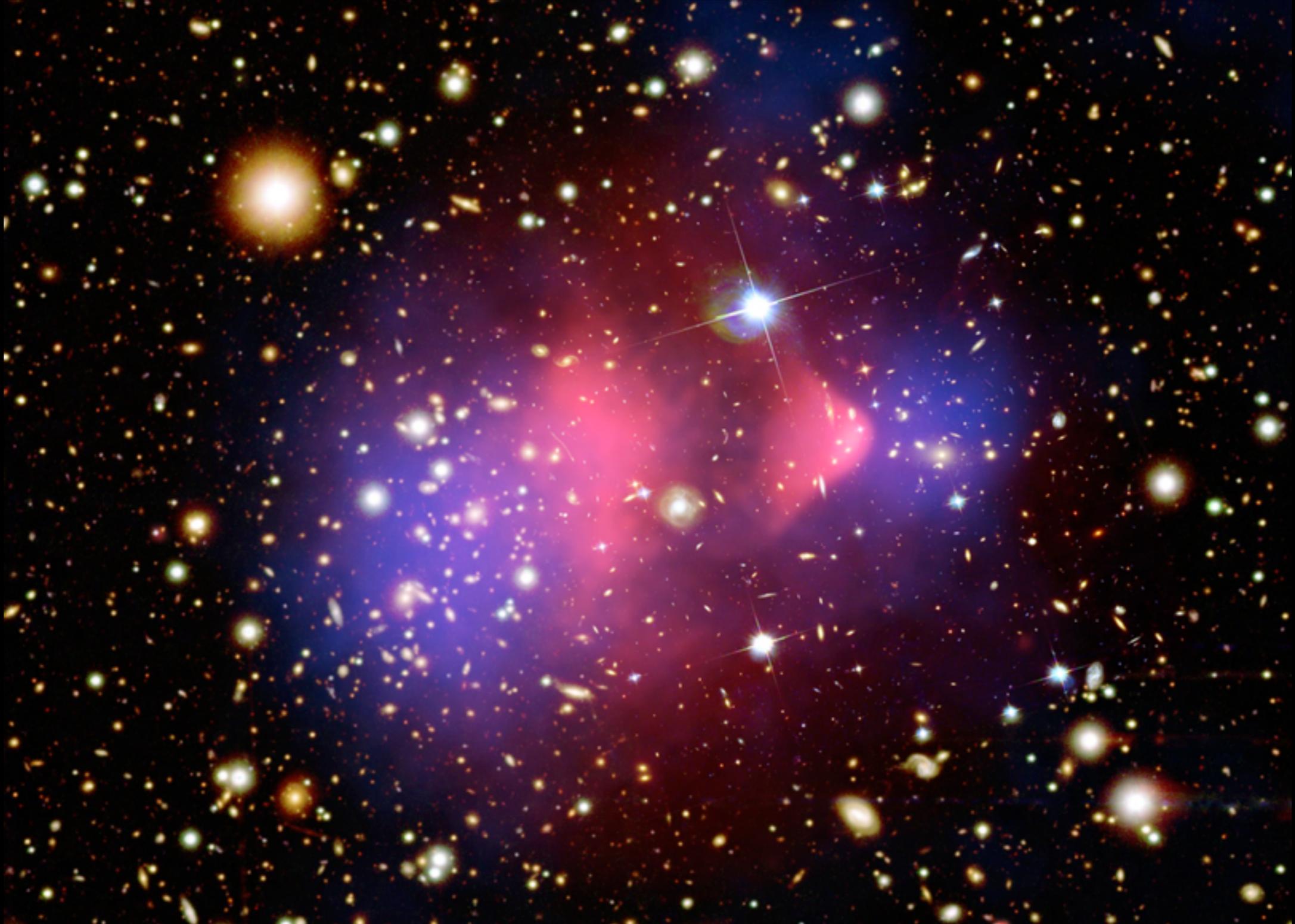
We've Rounded Up the Usual Suspects.....

**But sometimes science takes us in
unexpected directions**

But first we should double check our measurements.

How do we know that the odd rotations of galaxies are real evidence for a new type of matter?

Bullet Cluster

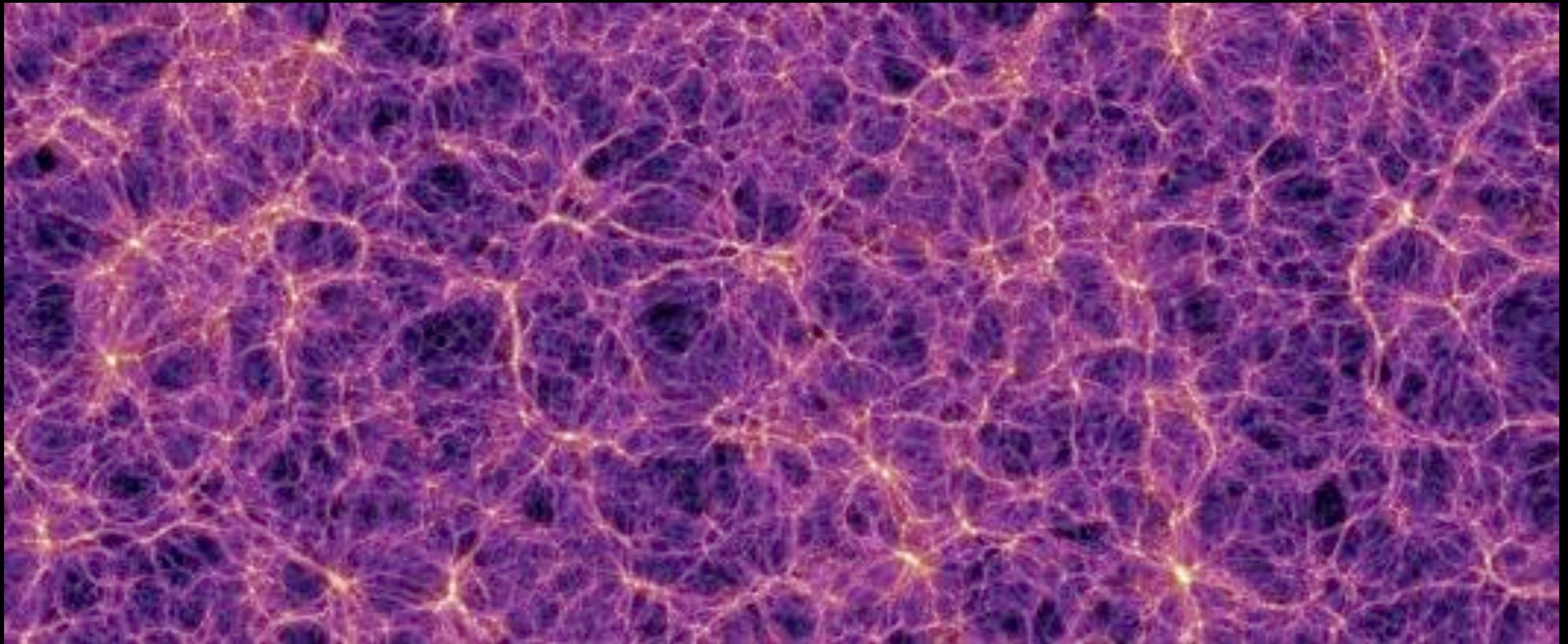


Formation of Structure



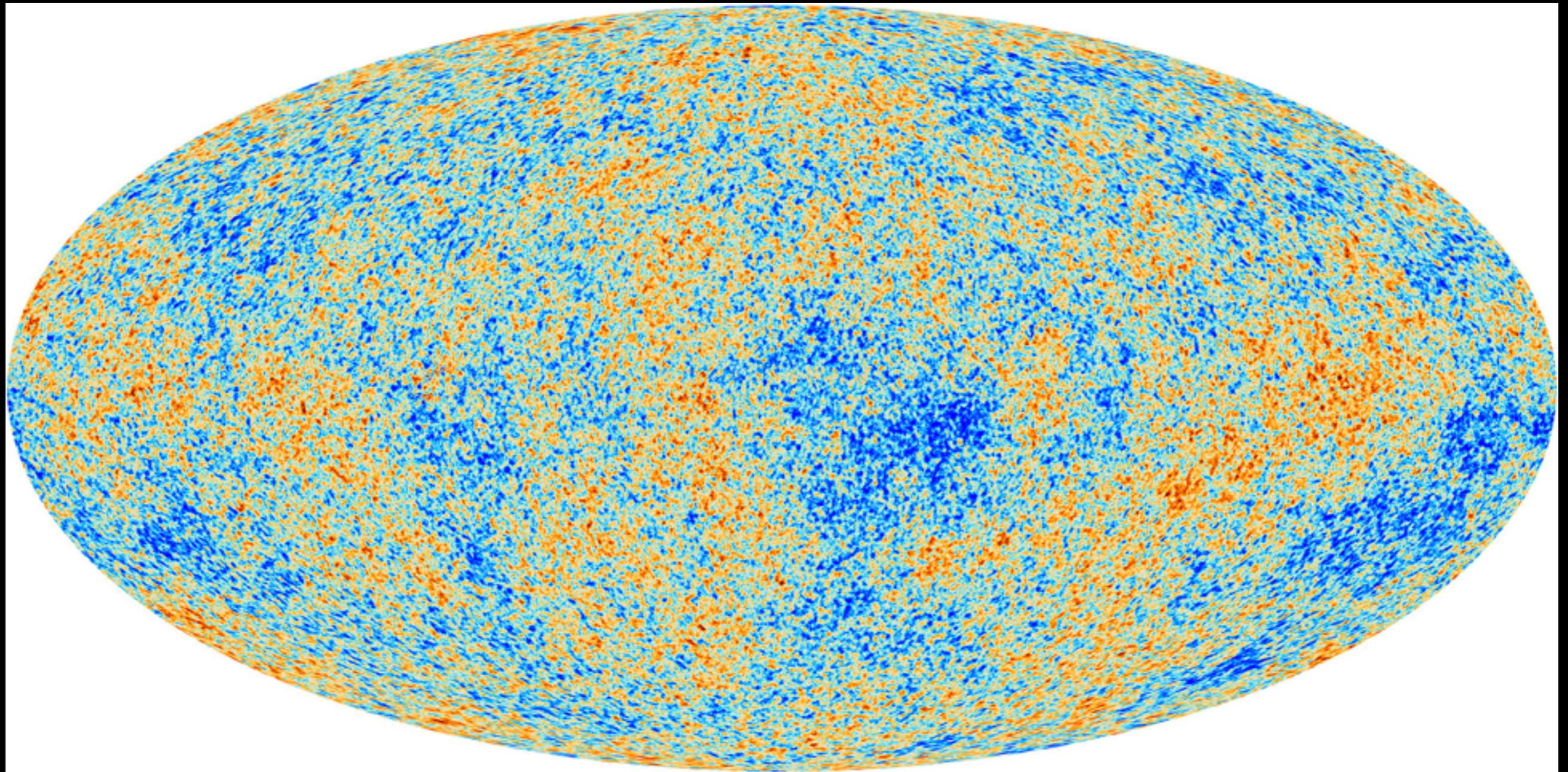
We can produce computer simulations of how the galaxies in the universe form, when we either add in, or leave out, a new form of matter that doesn't interact with light.

Formation of Structure

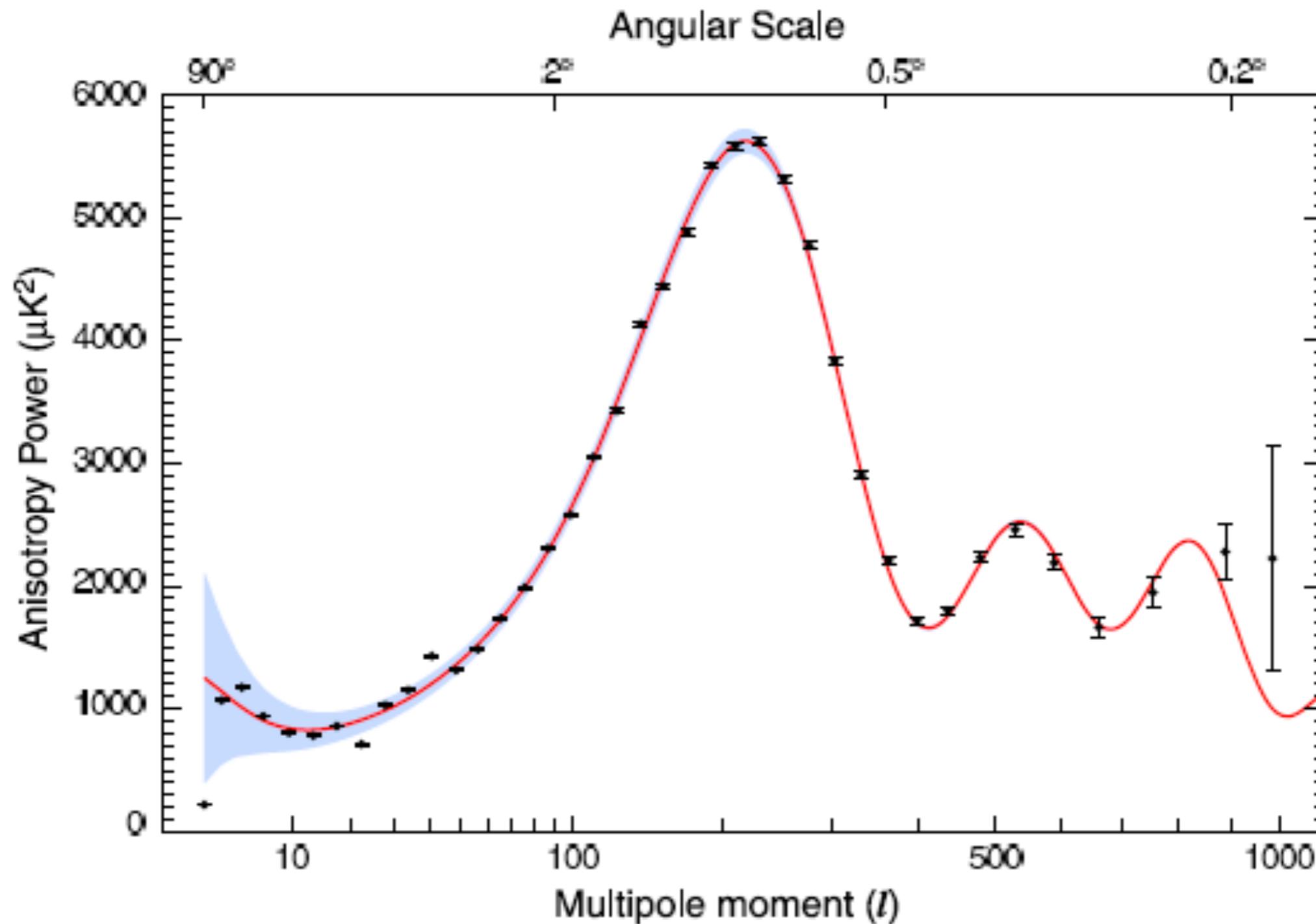


The simulations with dark matter match the observed distribution of galaxies, while simulations without dark matter do not.

Cosmic-Microwave Background



Cosmic-Microwave Background

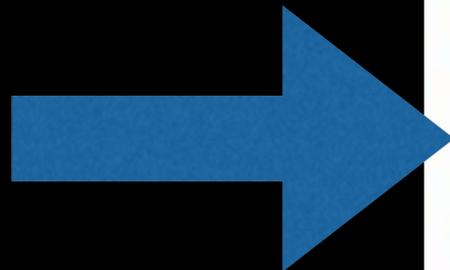


Particle Physics

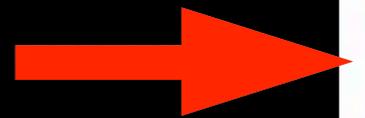
- We need to find a type of matter that can reproduce the observed characteristics of dark matter:
 - **1.) Dark** - Doesn't interact with photon
 - **2.) Stable** - Around at beginning of universe, and today
 - **3.) Slow-Moving (heavy)** - Each individual particle must be relatively heavy.

Standard Model

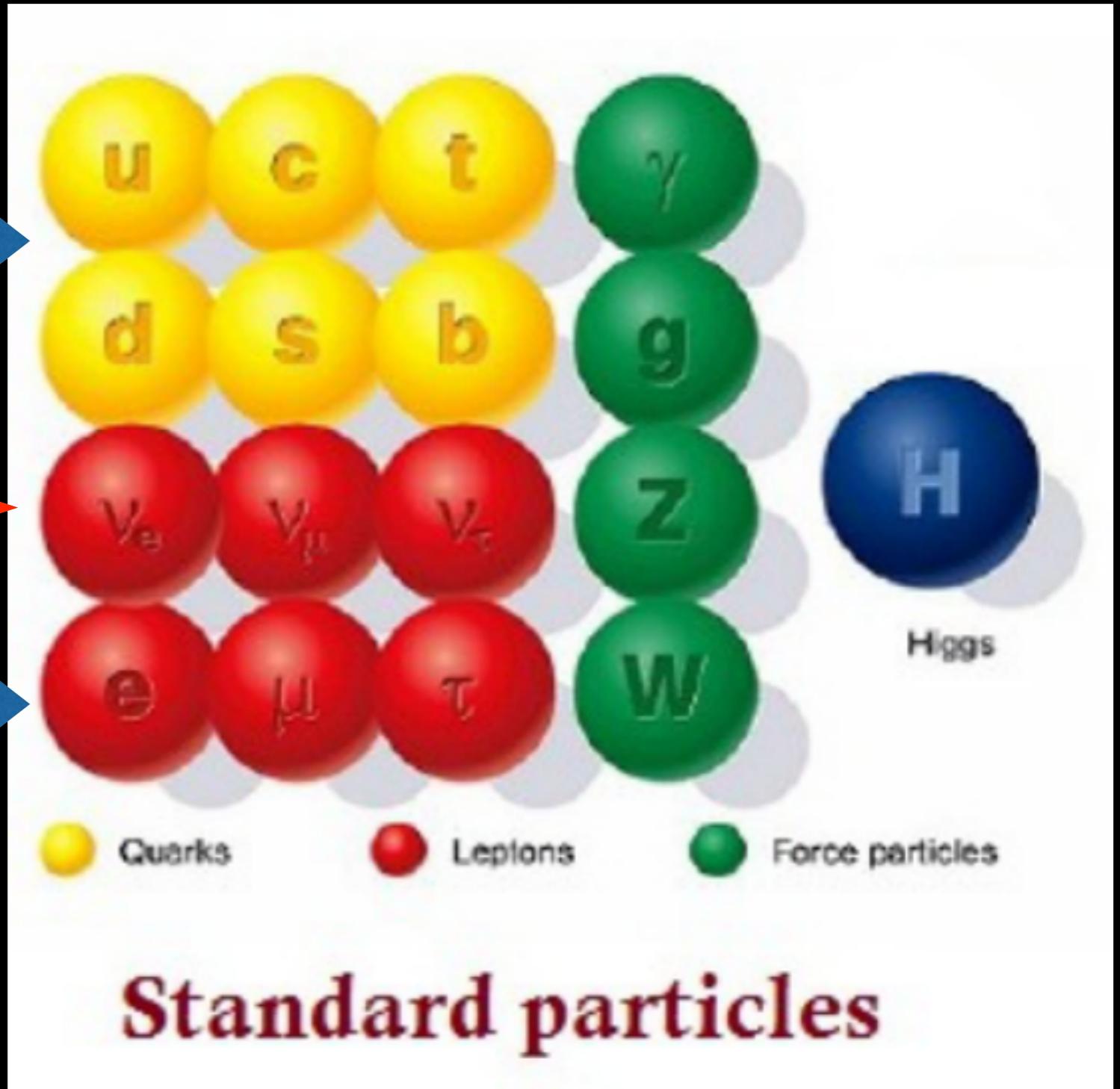
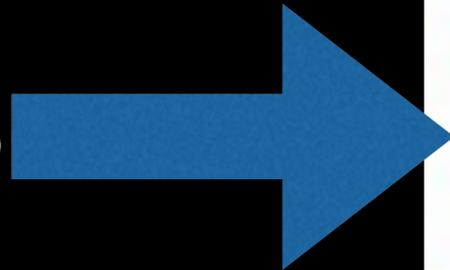
Protons



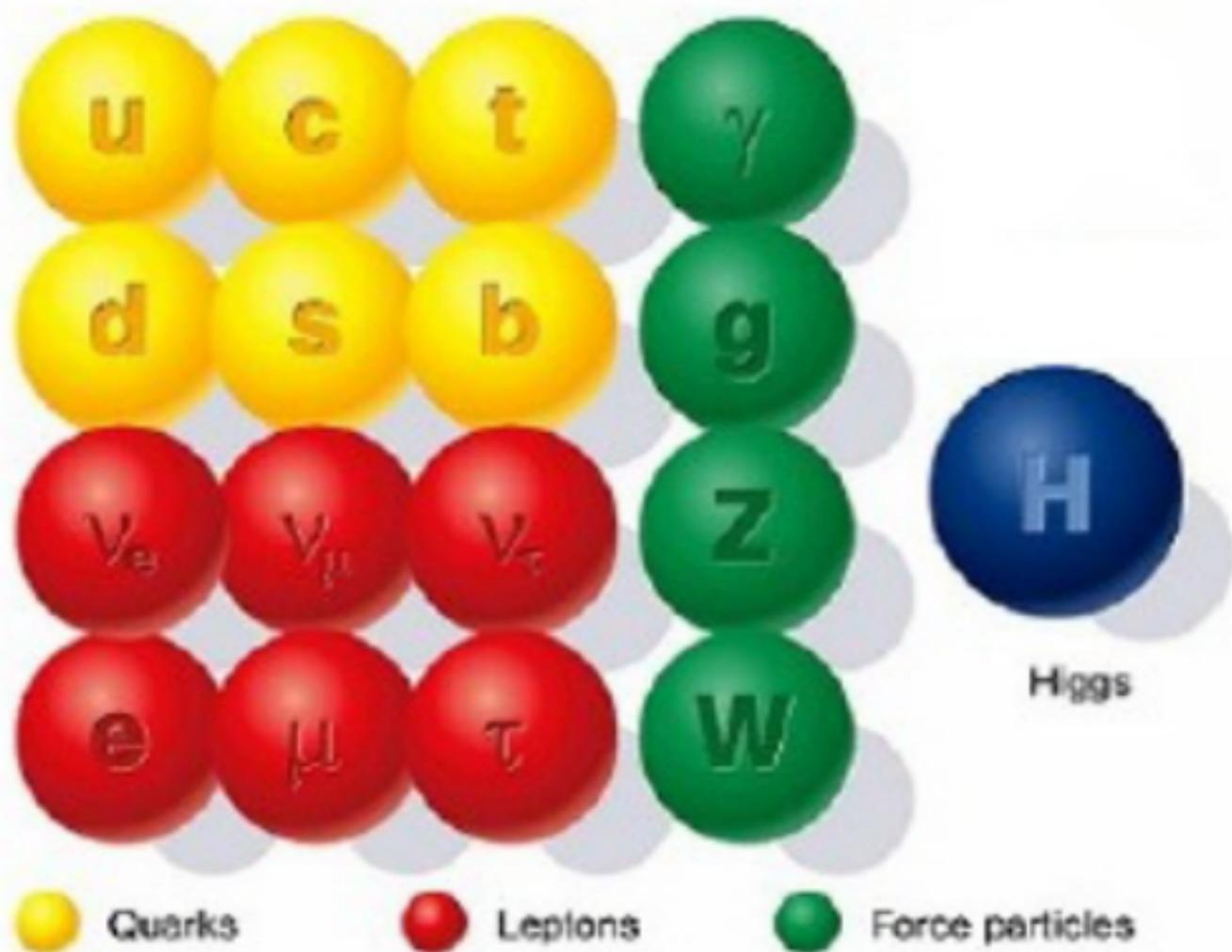
Neutrinos



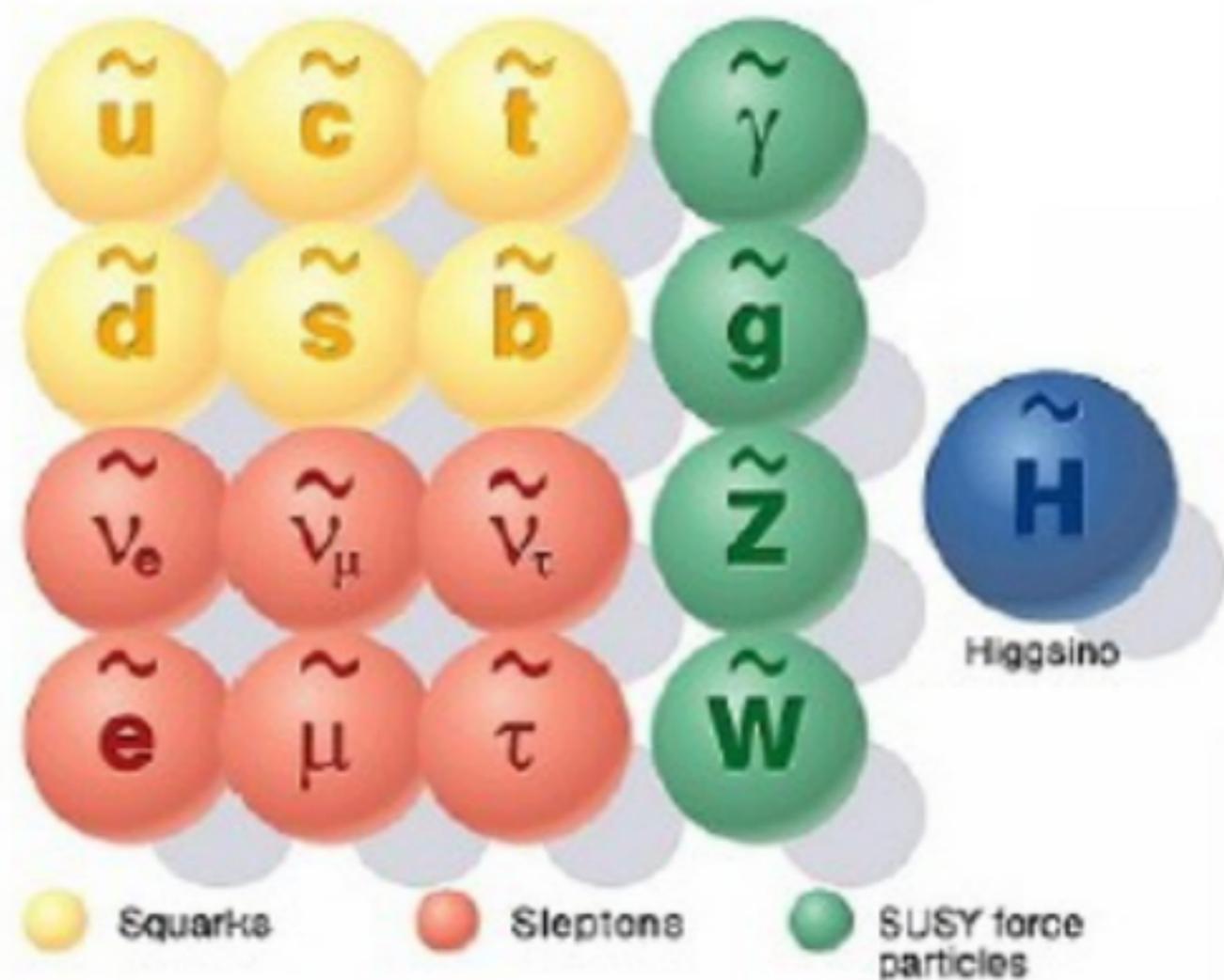
Electrons



SUPERSYMMETRY

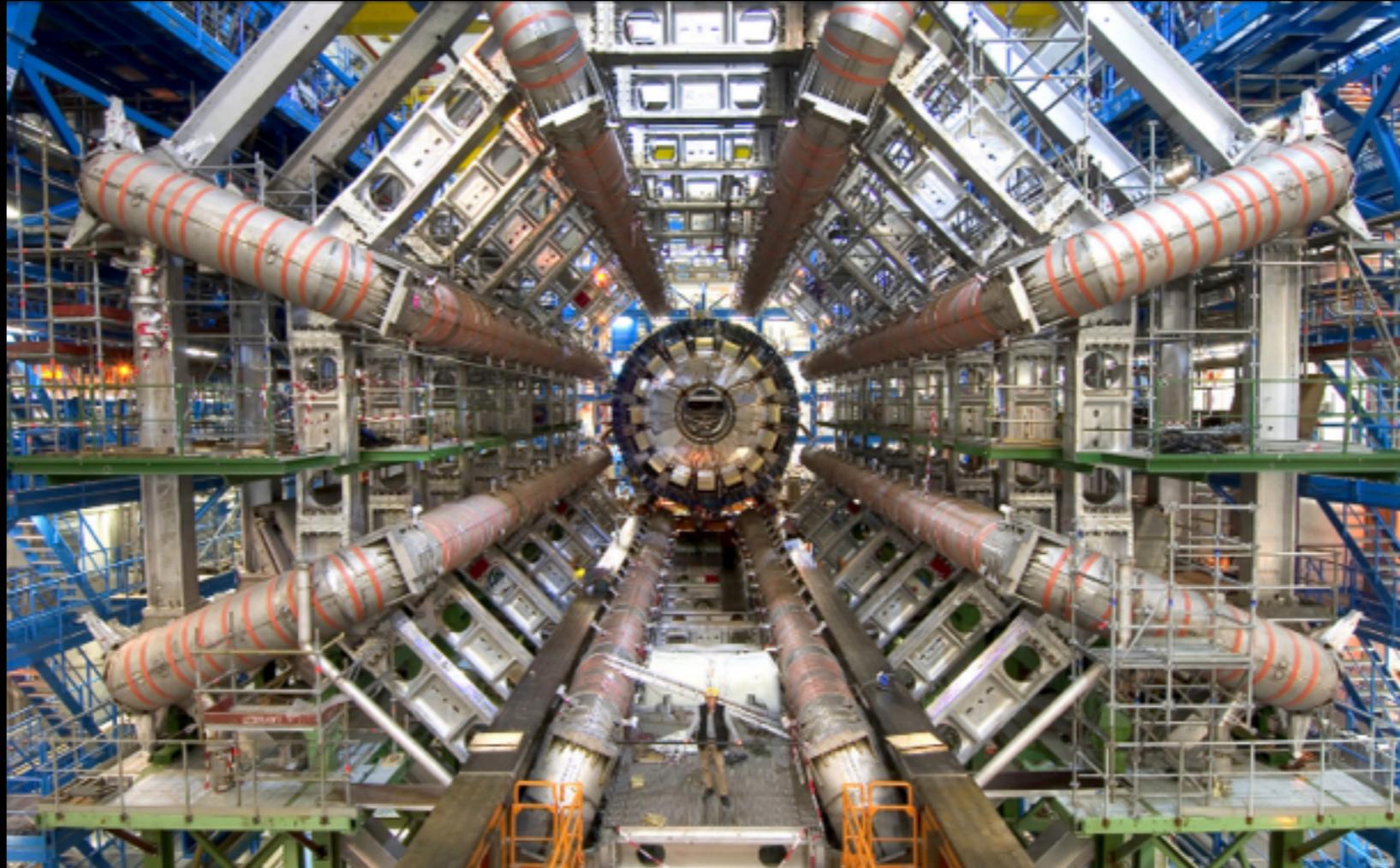


Standard particles



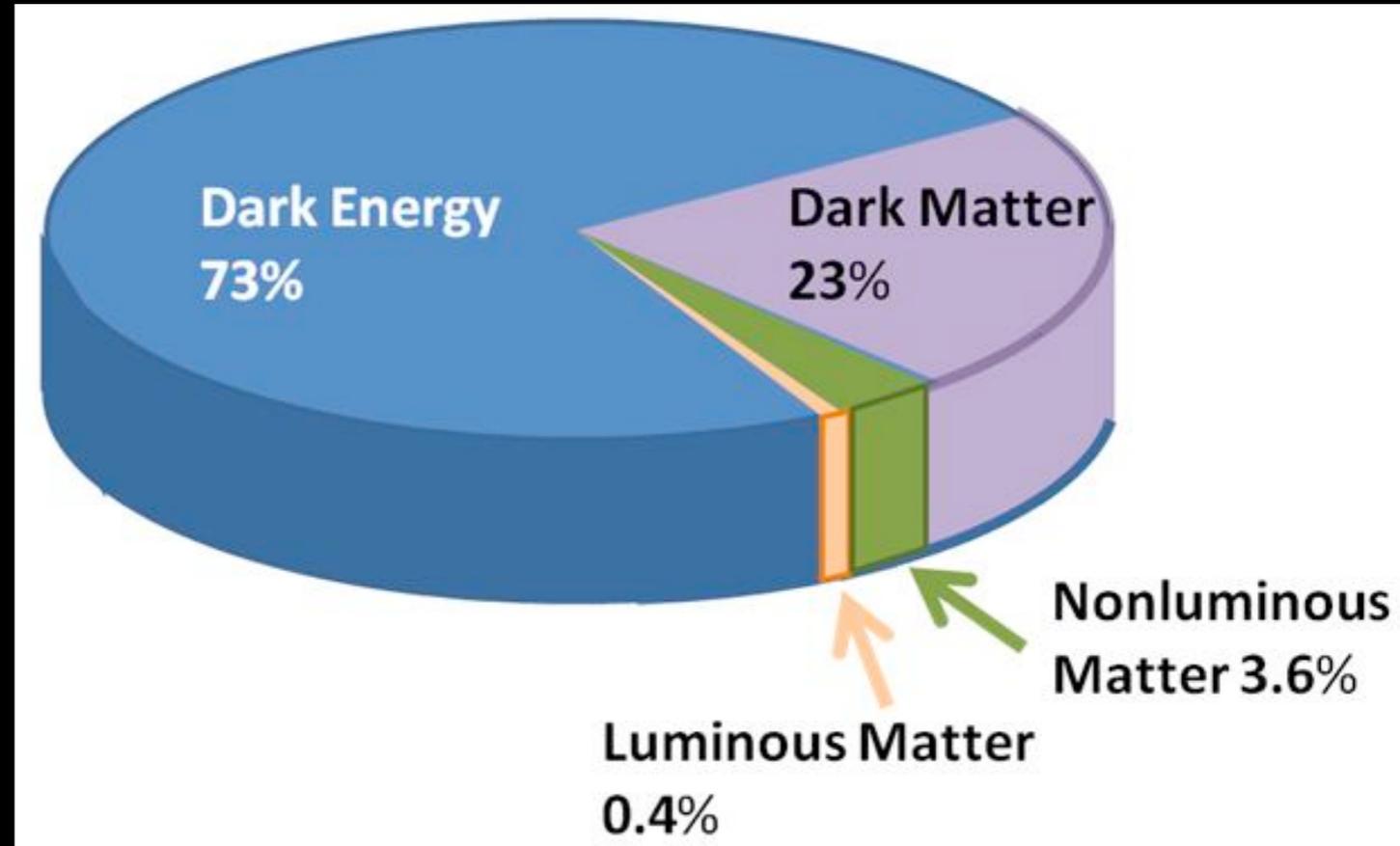
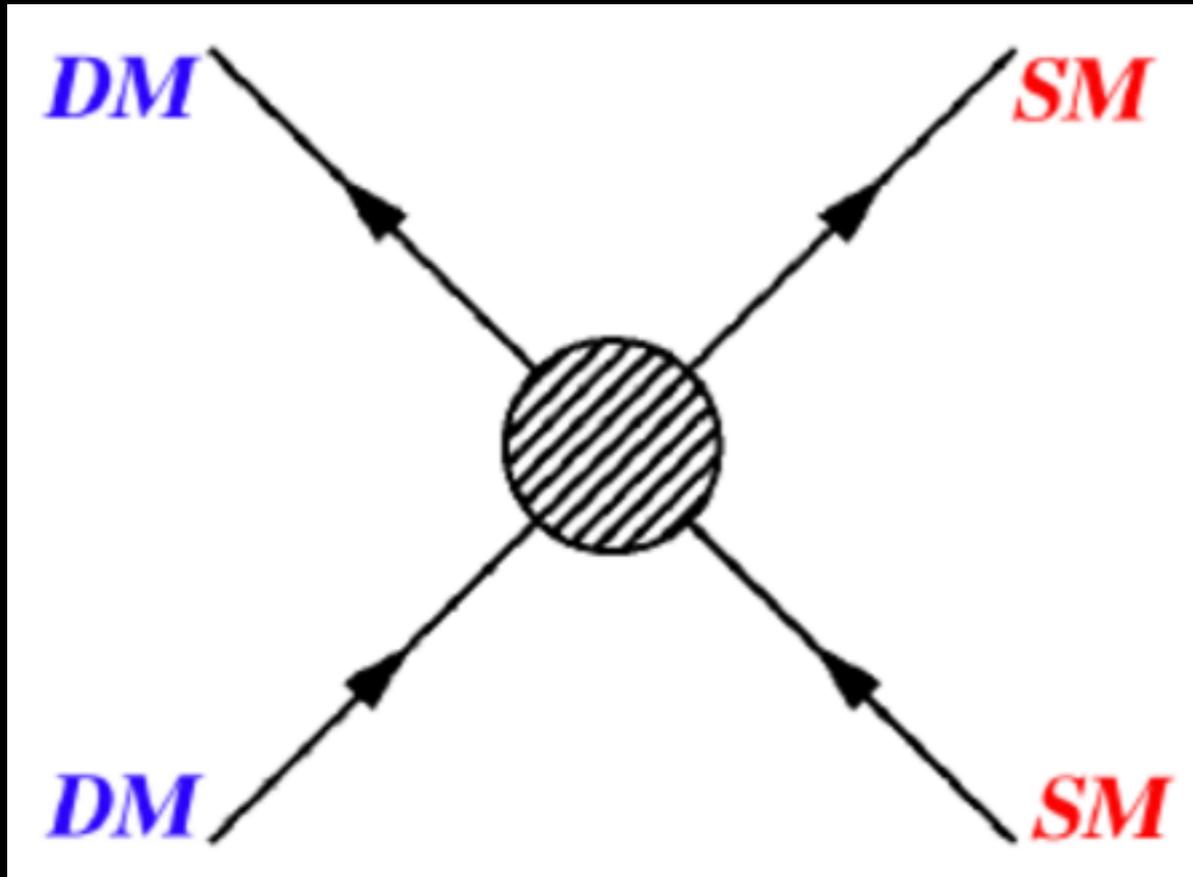
SUSY particles

Higgs Boson



- **Particle Accelerators recently succeeded in finding the Higgs Boson, which has important implications for the possible masses and interactions of supersymmetric particles**

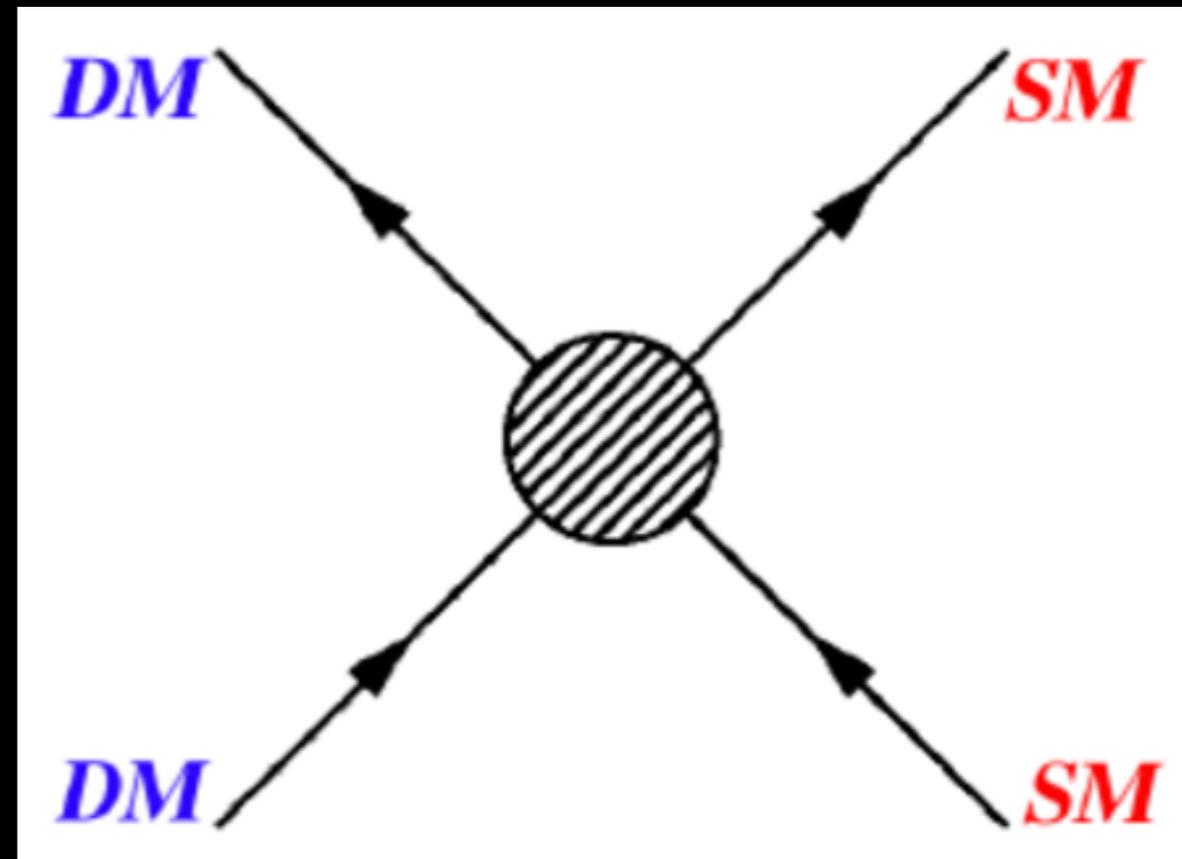
WIMP Miracle!



- In the late 1970s, it was discovered that if you gave allowed dark matter to interact via the weak force (but not interact with light), then dark matter naturally obtained the correct (observed) density in the universe.

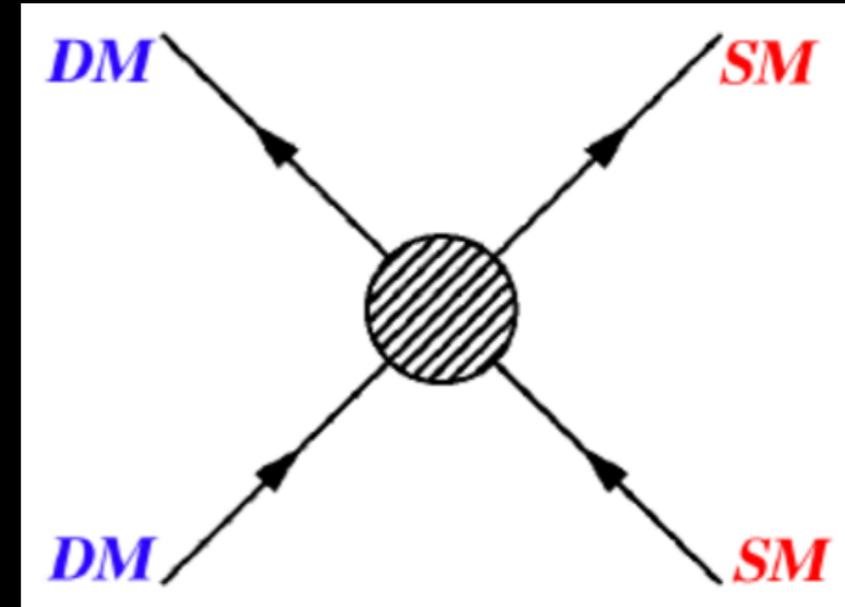
Detecting Dark Matter

- If dark matter interacts via the weak-force in the early universe, it should still interact via the weak force today.
- This means that we can search for interactions using this known force, in order to detect dark matter.



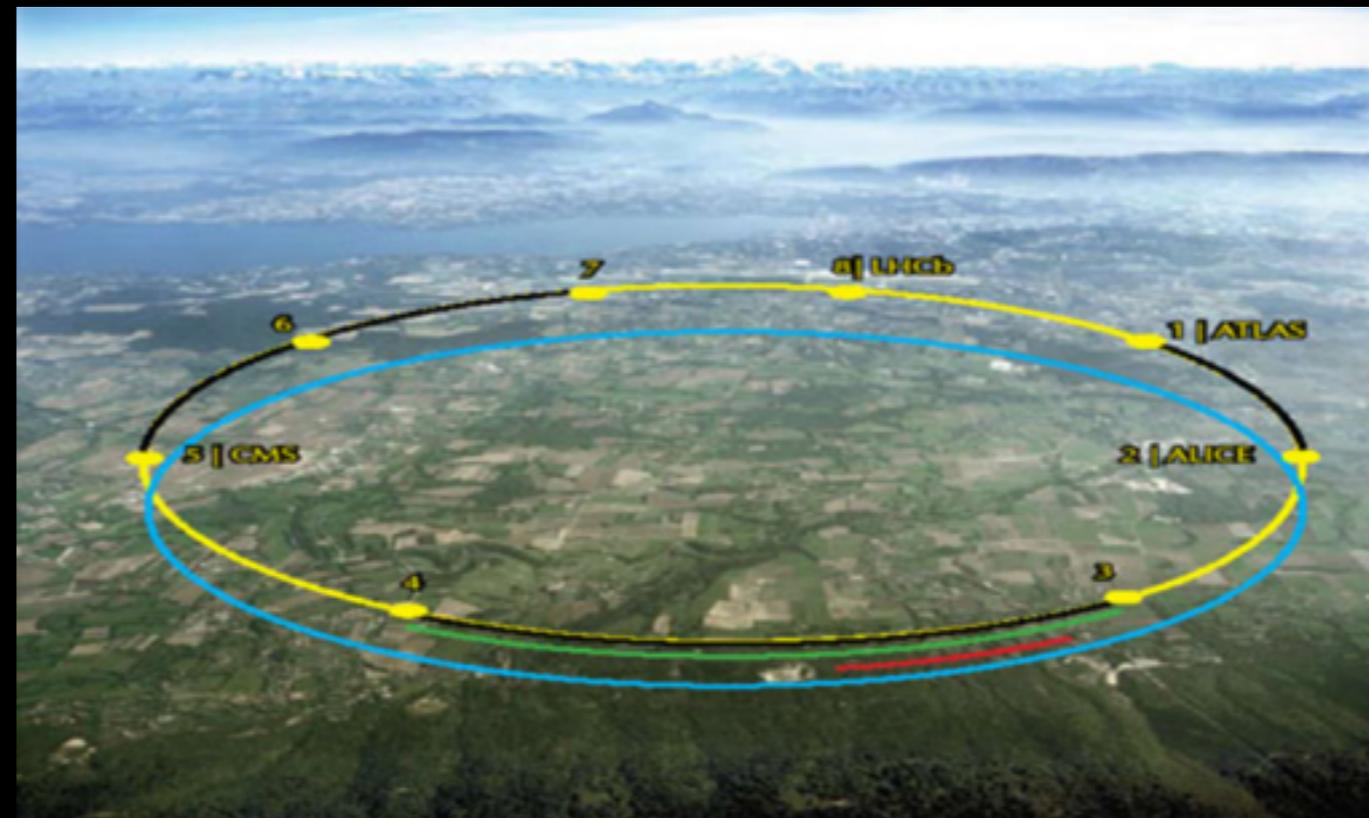
Particle Accelerators

- We can search for two normal (standard model) particles colliding, and producing two dark matter particles via this interaction.



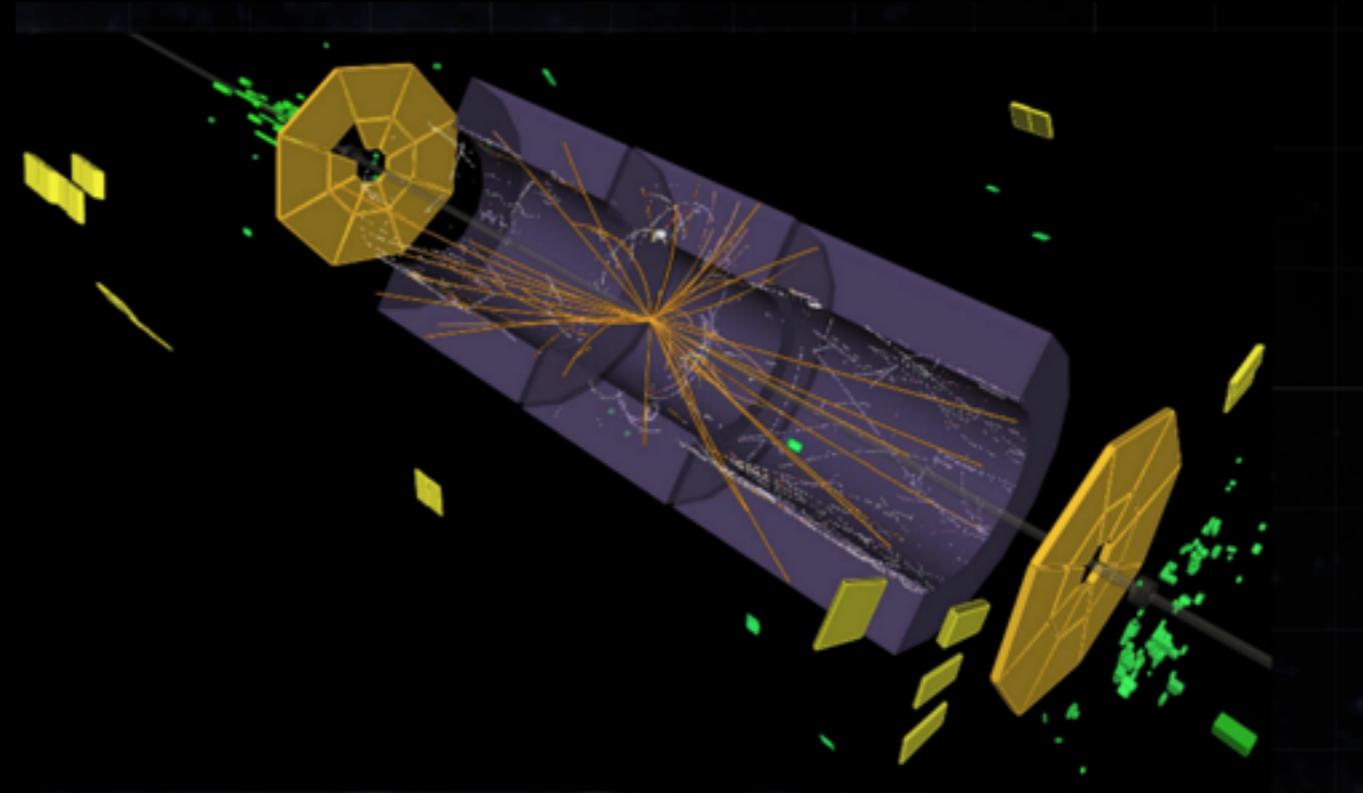
$$E = mc^2$$

- Since dark matter is heavy, we must accelerate the standard model particles to very high energies, in order to produce the dark matter particles



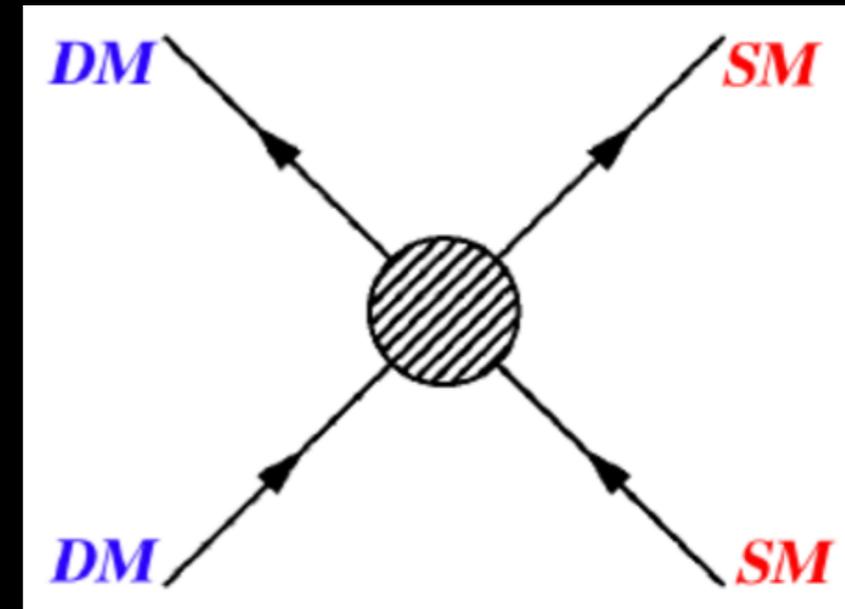
Problem! Dark Matter is Hard to See

- So we've made the dark matter particle.
- But now we need to see it with our instrument.
- Instead we can look for missing energy, in collisions that produce particles we can see.

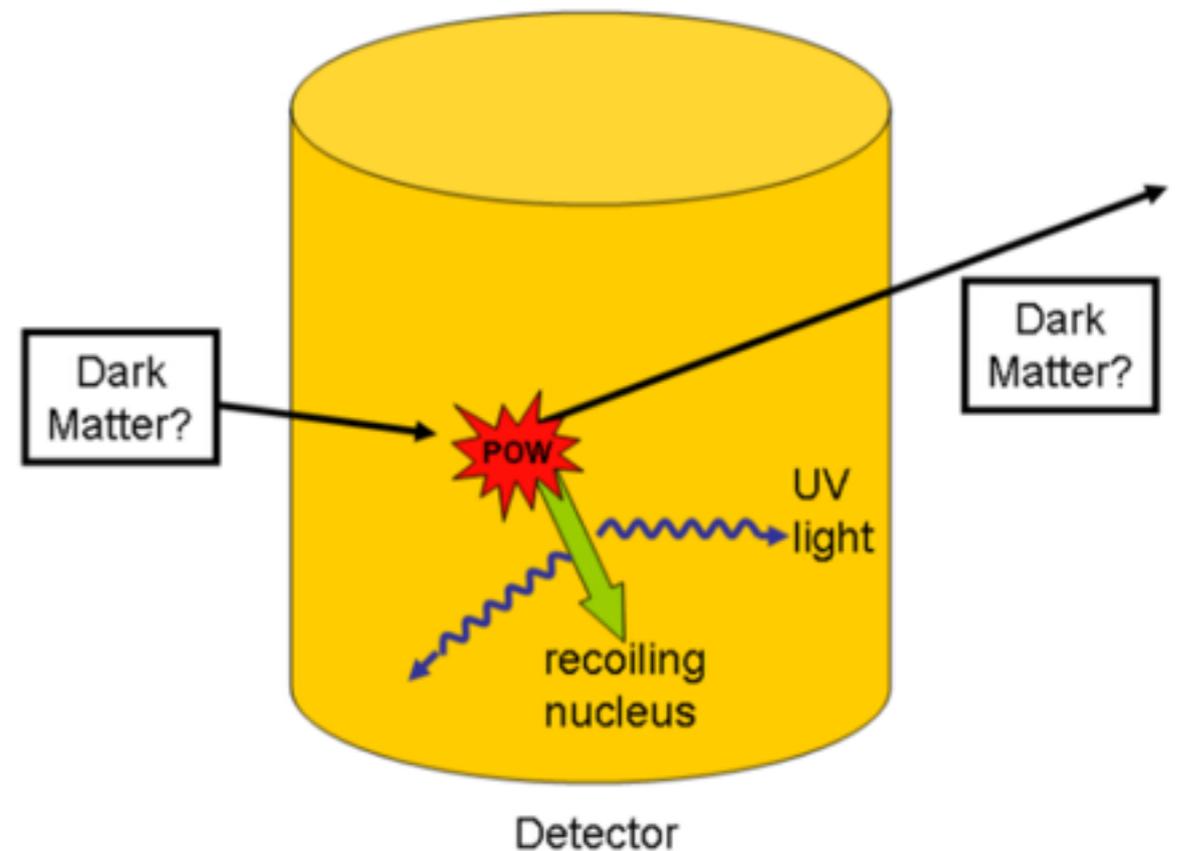


Direct Detection

- Dark Matter is all around us - traveling through us at this very second.



- We can look for the rare weak force interactions between dark matter and normal matter.



Direct Detection

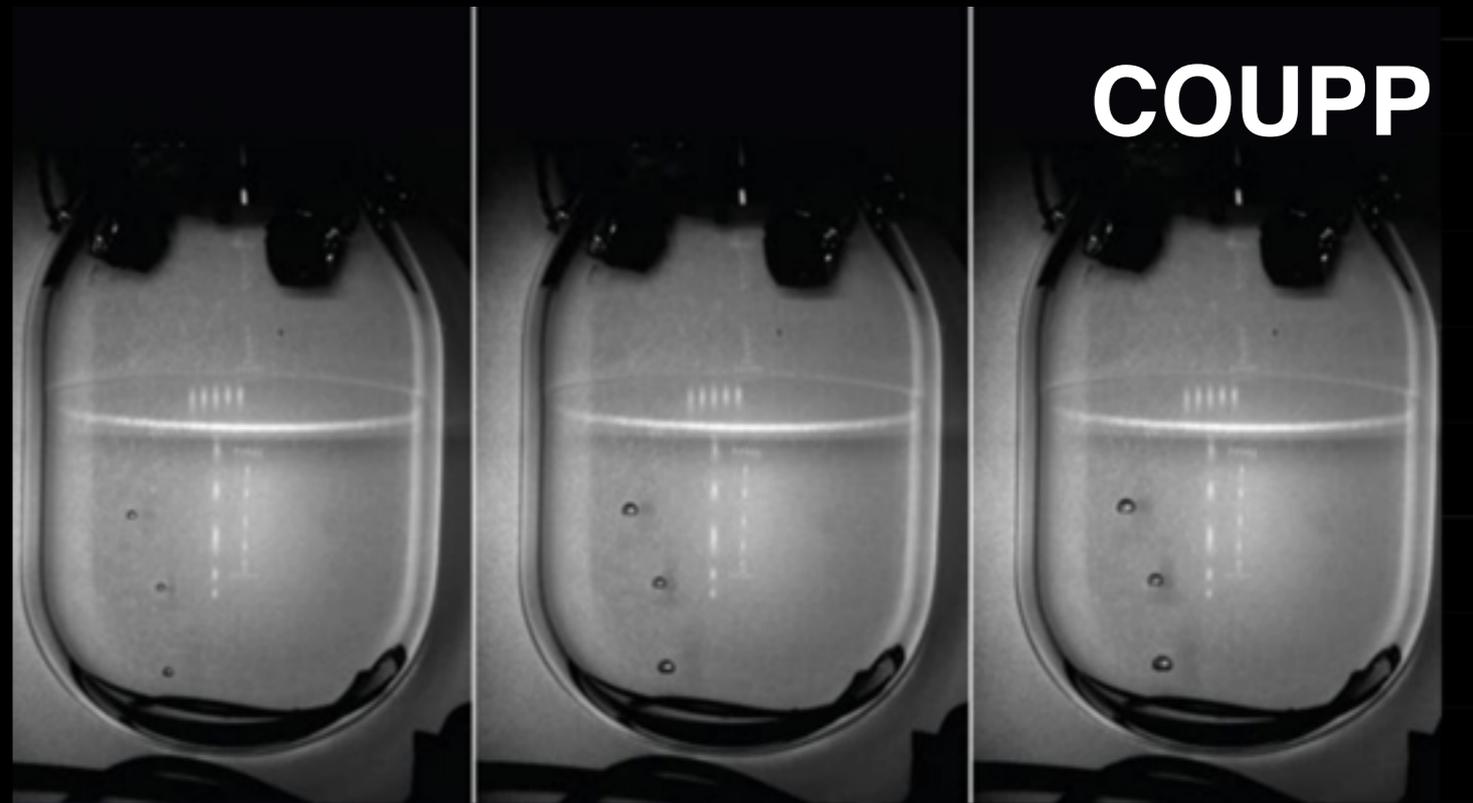
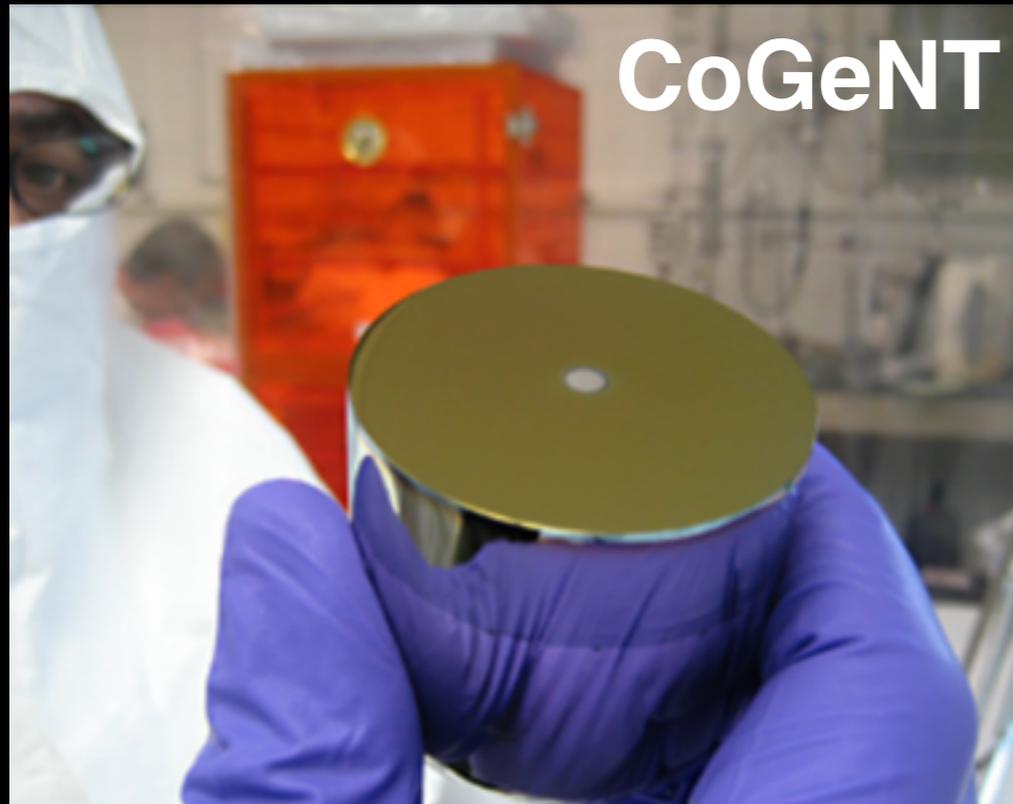
- **Problem! We need a way to distinguish the “jiggles” due to dark matter interactions from dark matter interactions due to normal particles**
- **Need to go deep underground to block the normal particles.**



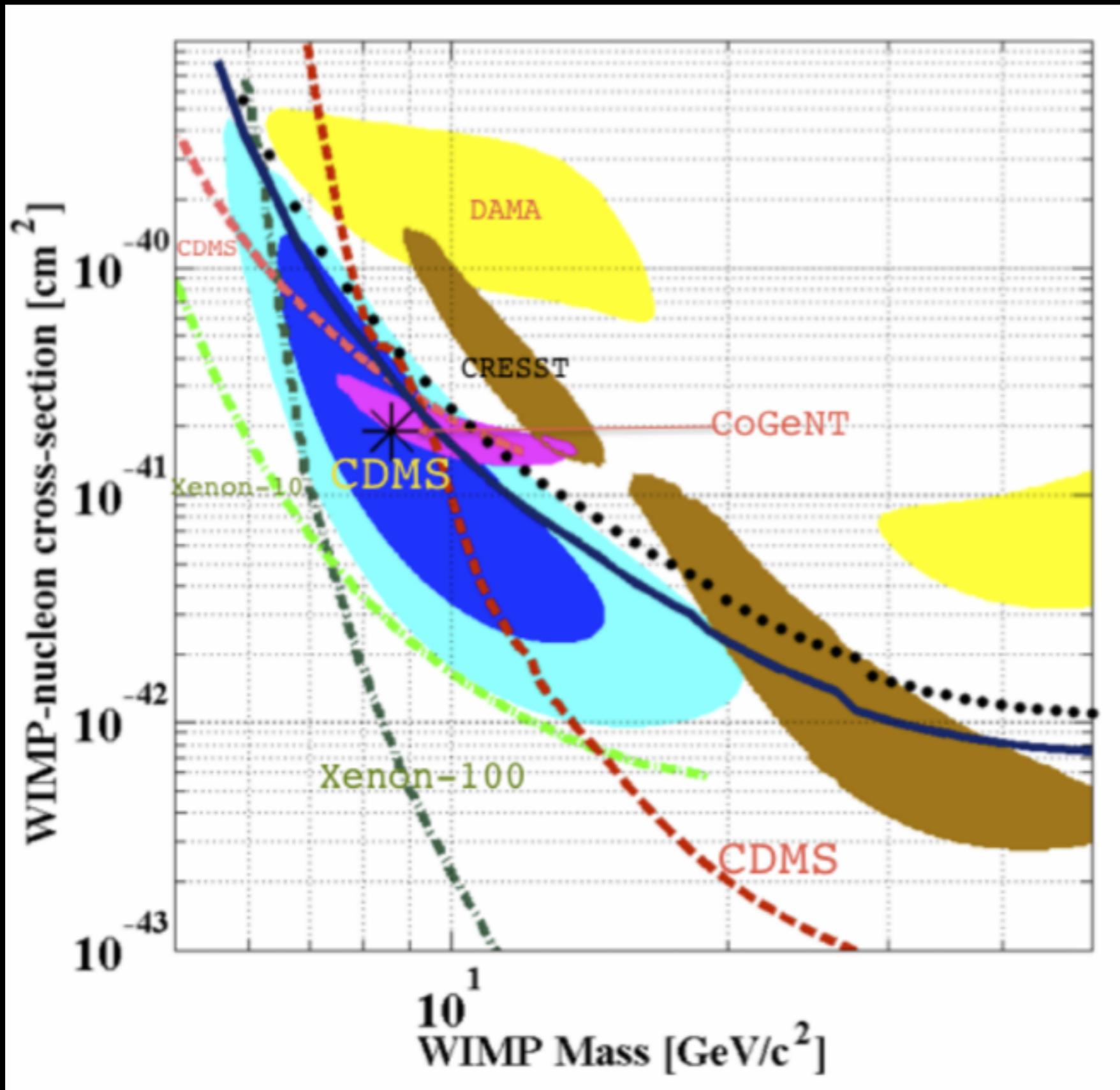
Direct Detection



Chicago Connections

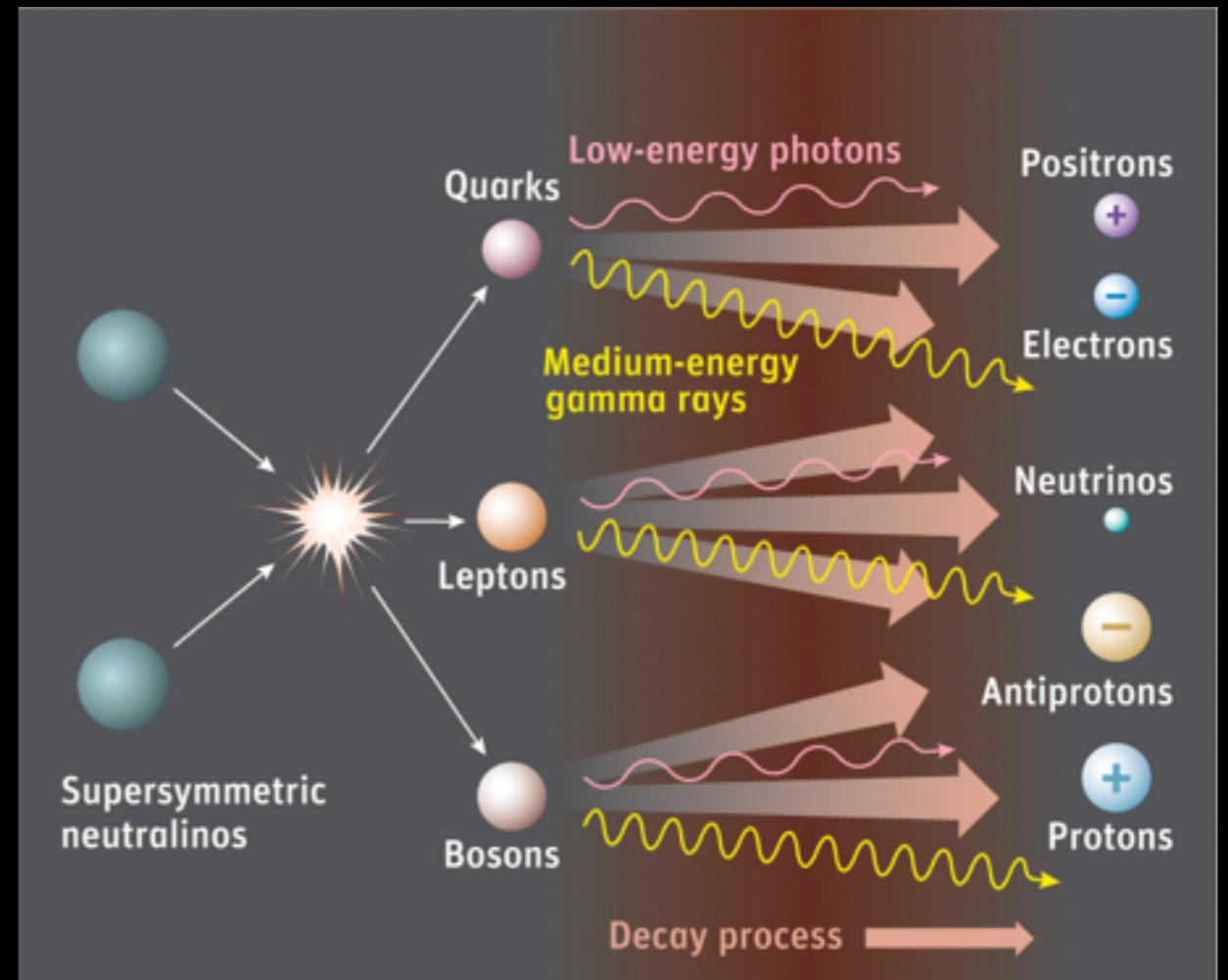
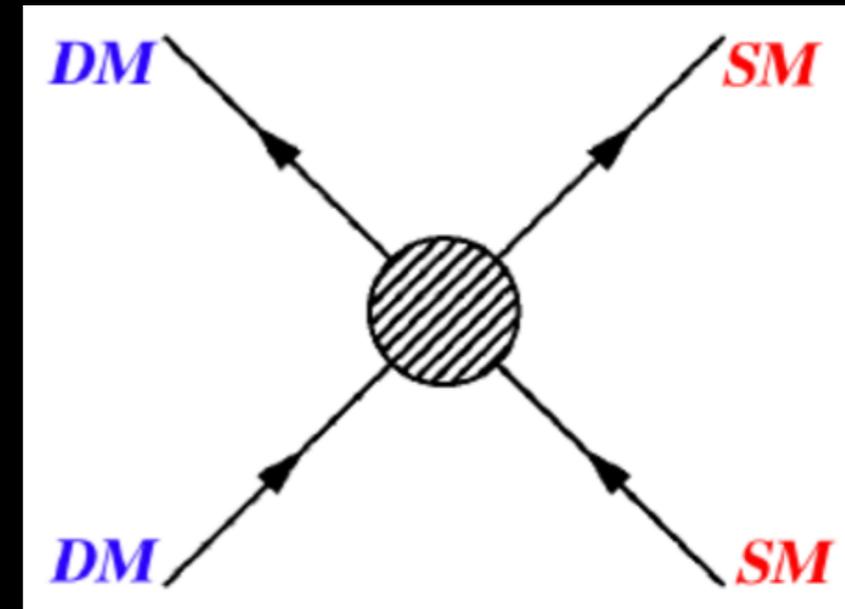


Results Are A Mess

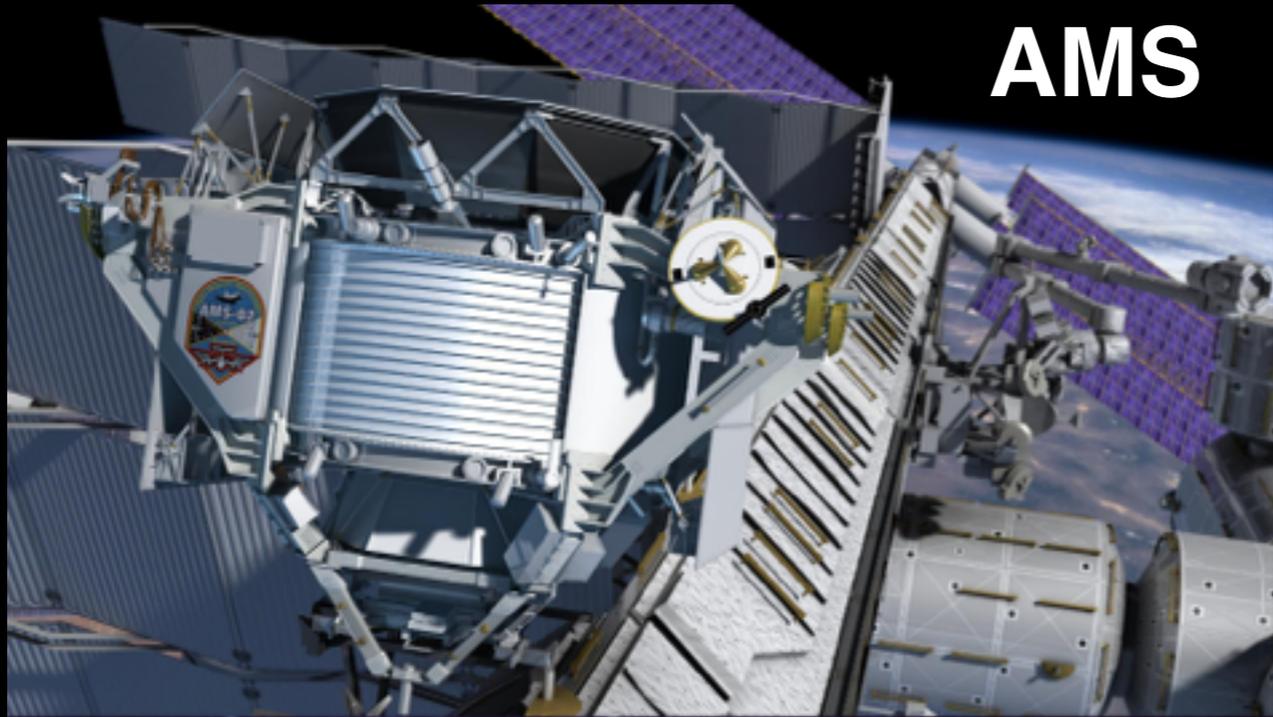


Indirect Detection

- Can also look for two dark matter particles interacting and producing standard model particles
- We want to search in regions where there is lots of dark matter - go to space!



Indirect Detection



AMS



Fermi-LAT



VERITAS

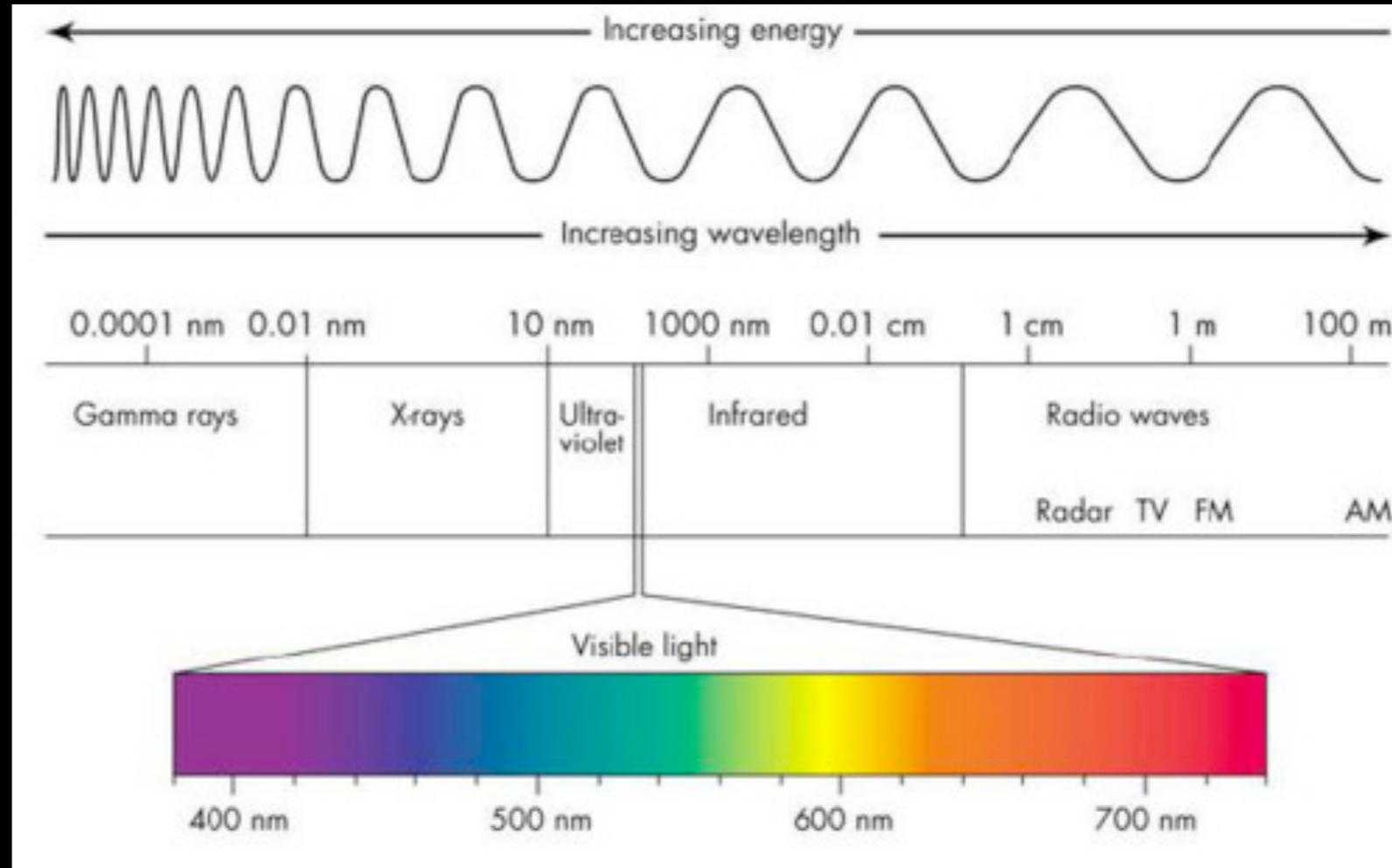


Indirect Detection

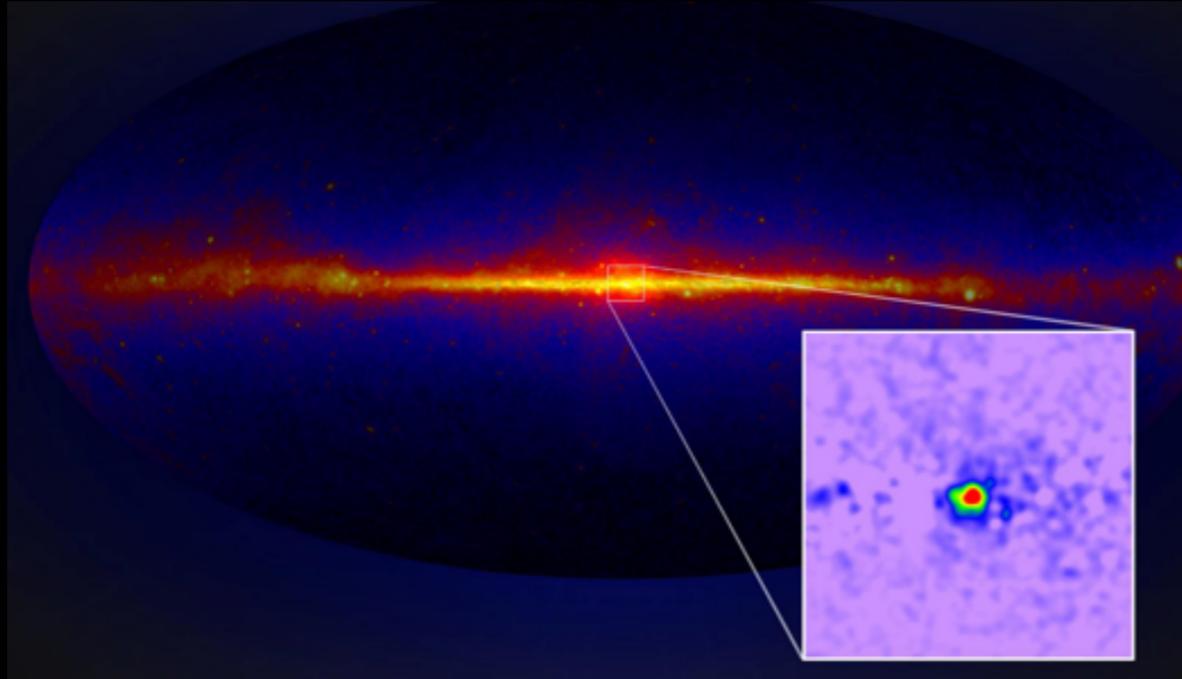
- What do these instruments have in common?

$$E = mc^2$$

- They look for very high-energy standard model particles.



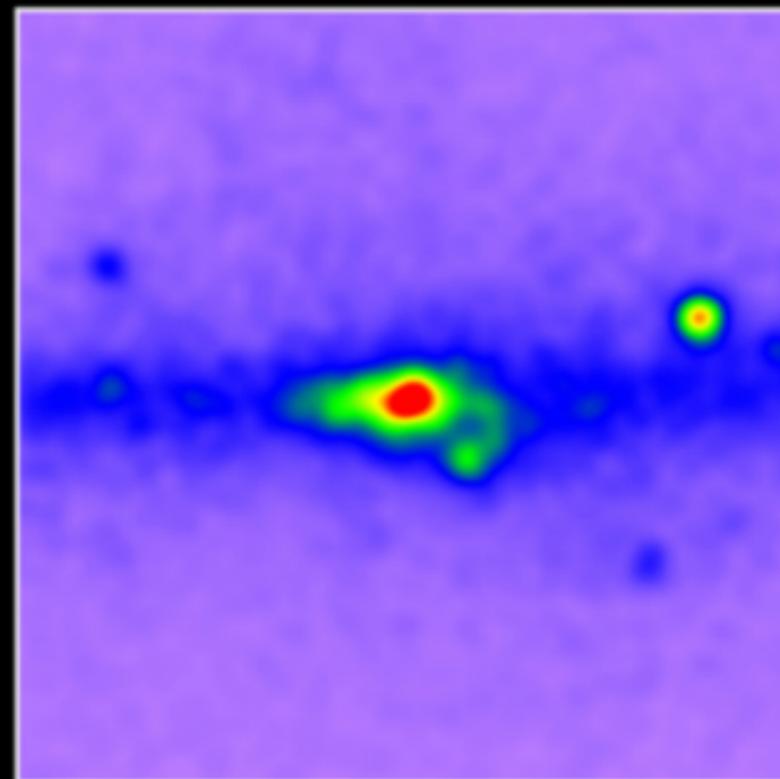
Indirect Detection



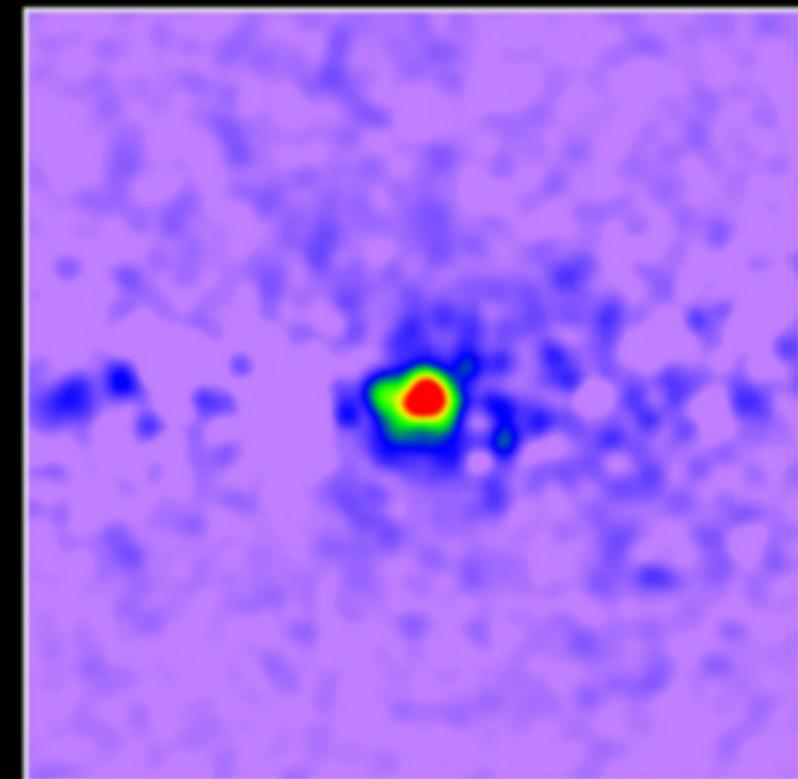
- We can look at the **Galactic Center**, which is one of the brightest sources for high energy light

Uncovering a gamma-ray excess at the galactic center

An excess is found, which contributes almost 1/3 of the total emission from the Galactic Center!



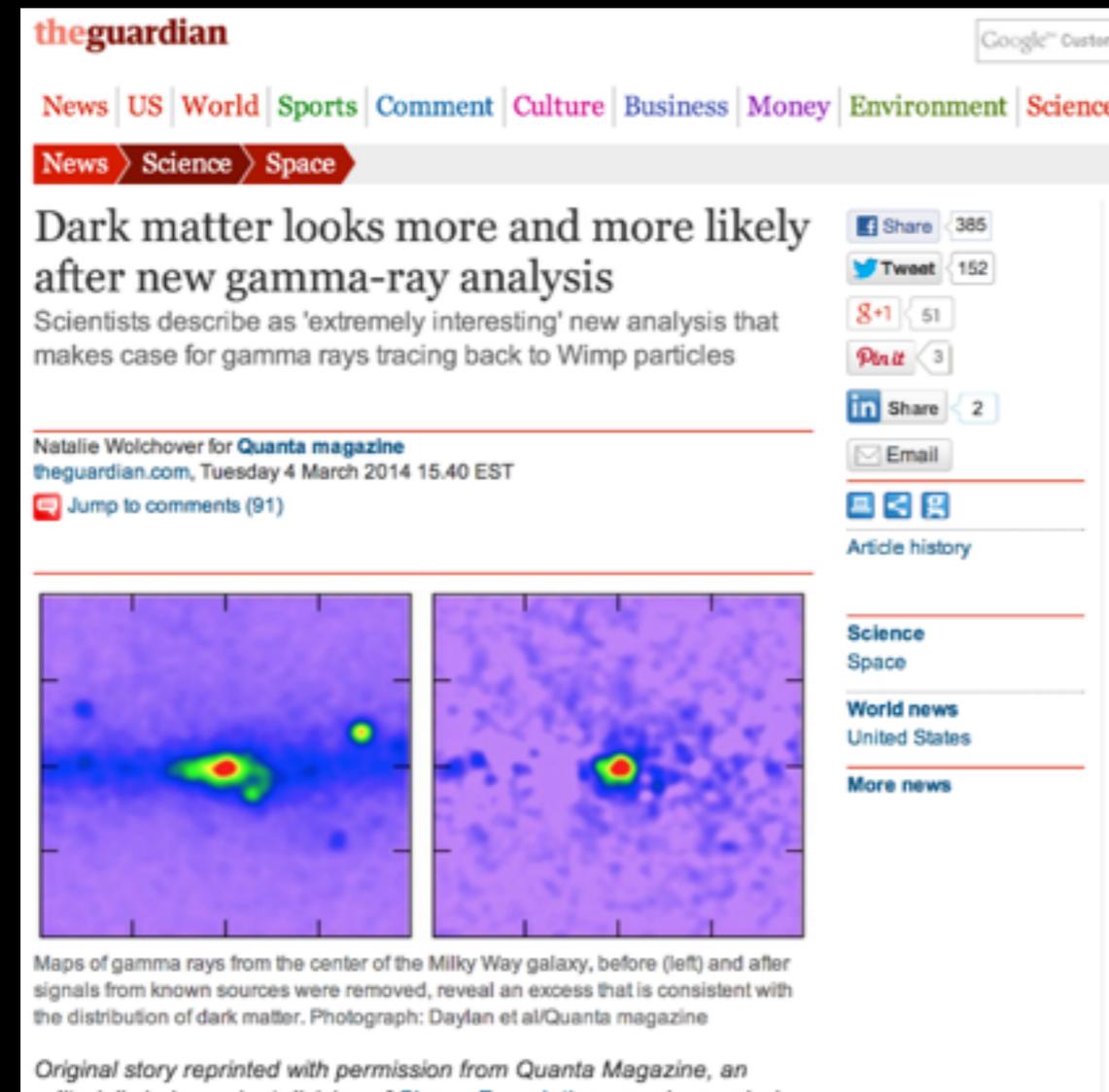
Unprocessed map of 1.0 to 3.16 GeV gamma rays



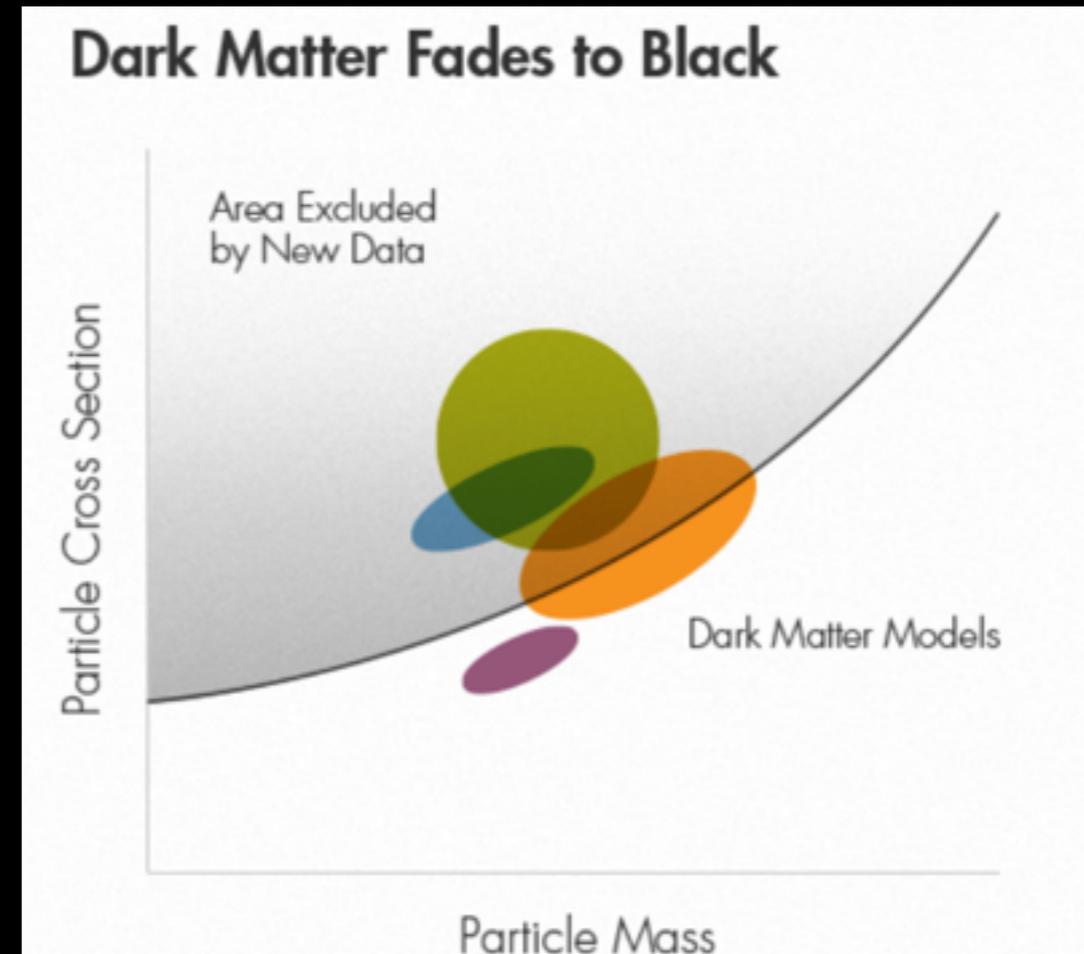
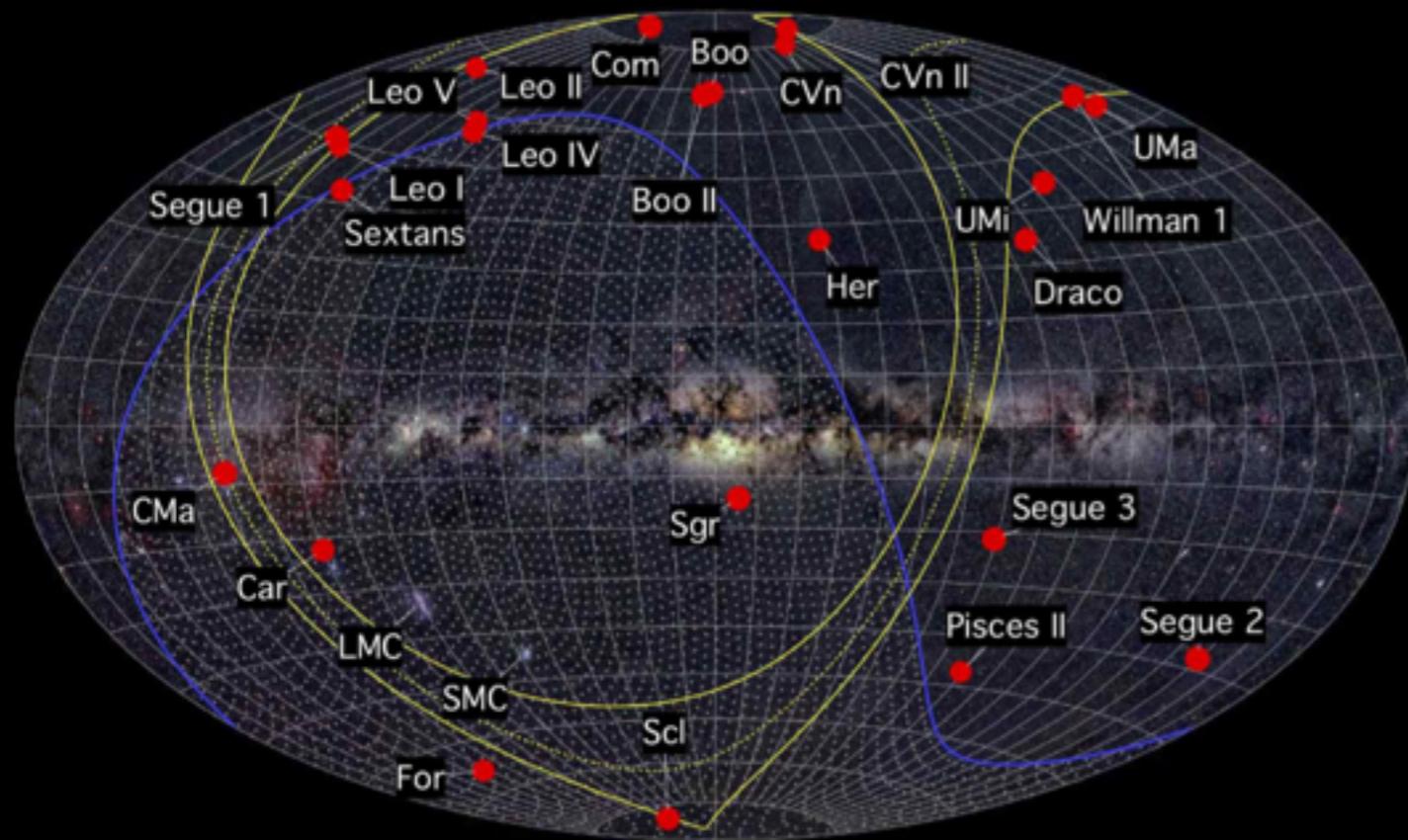
Known sources removed

Excitement!

This launched follow up investigations by many fellow researchers

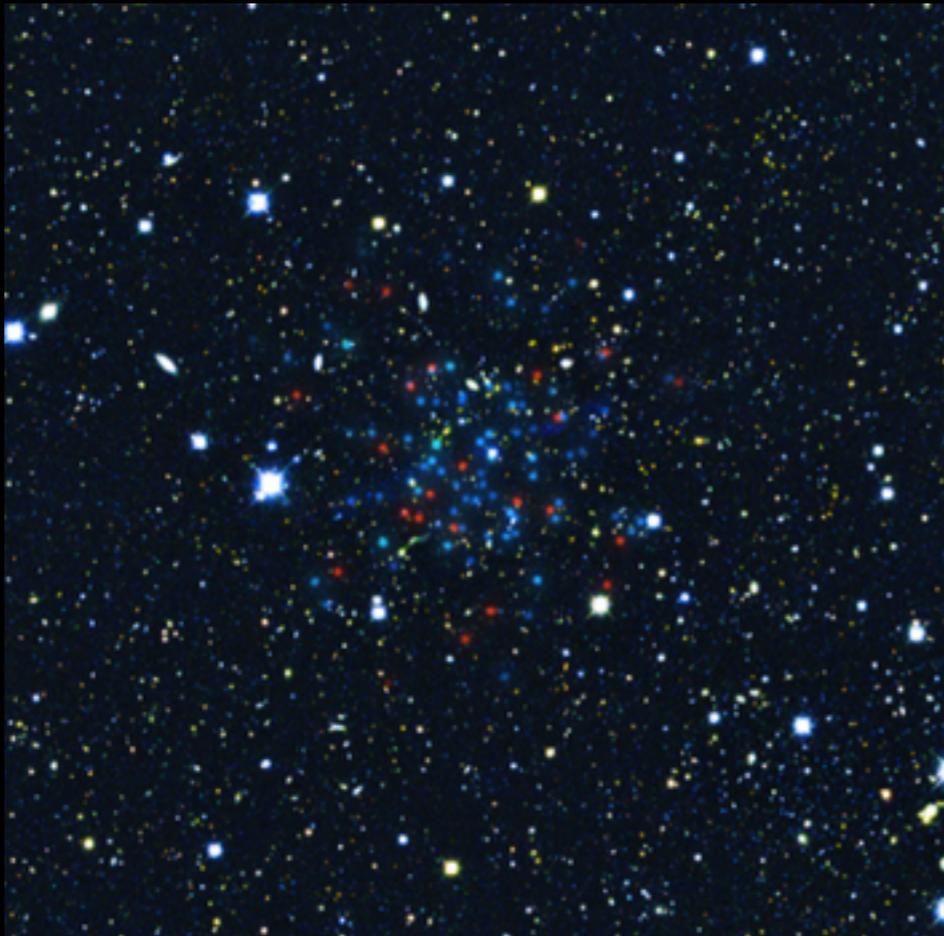


Dwarf Galaxies



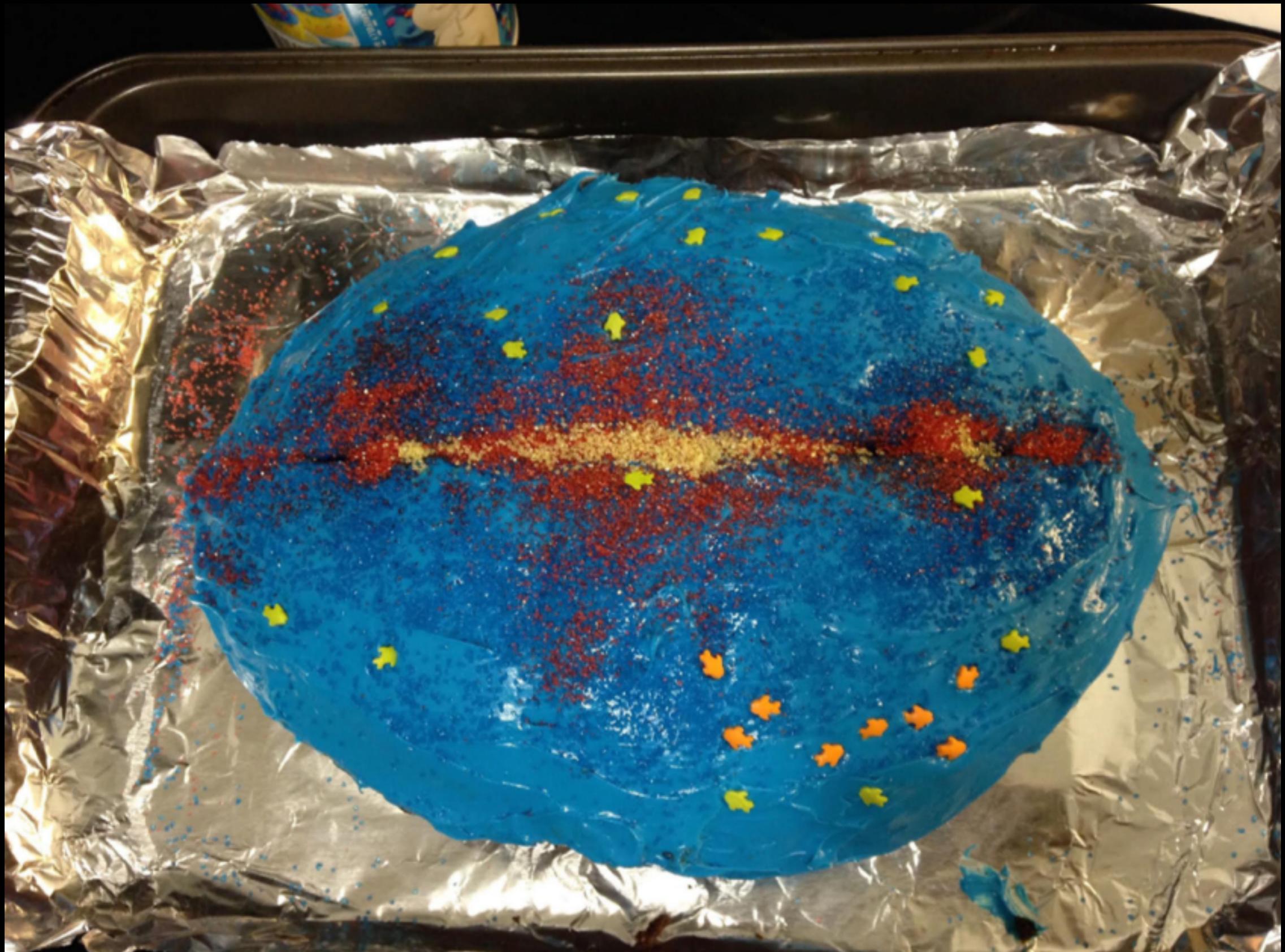
- **But if this is dark matter, we want to find the same signal everywhere we look in the sky. The second brightest signal should come from the population of minor “dwarf” galaxies around the Milky Way**
- **And no excess is found.**

Hot Off the Presses!

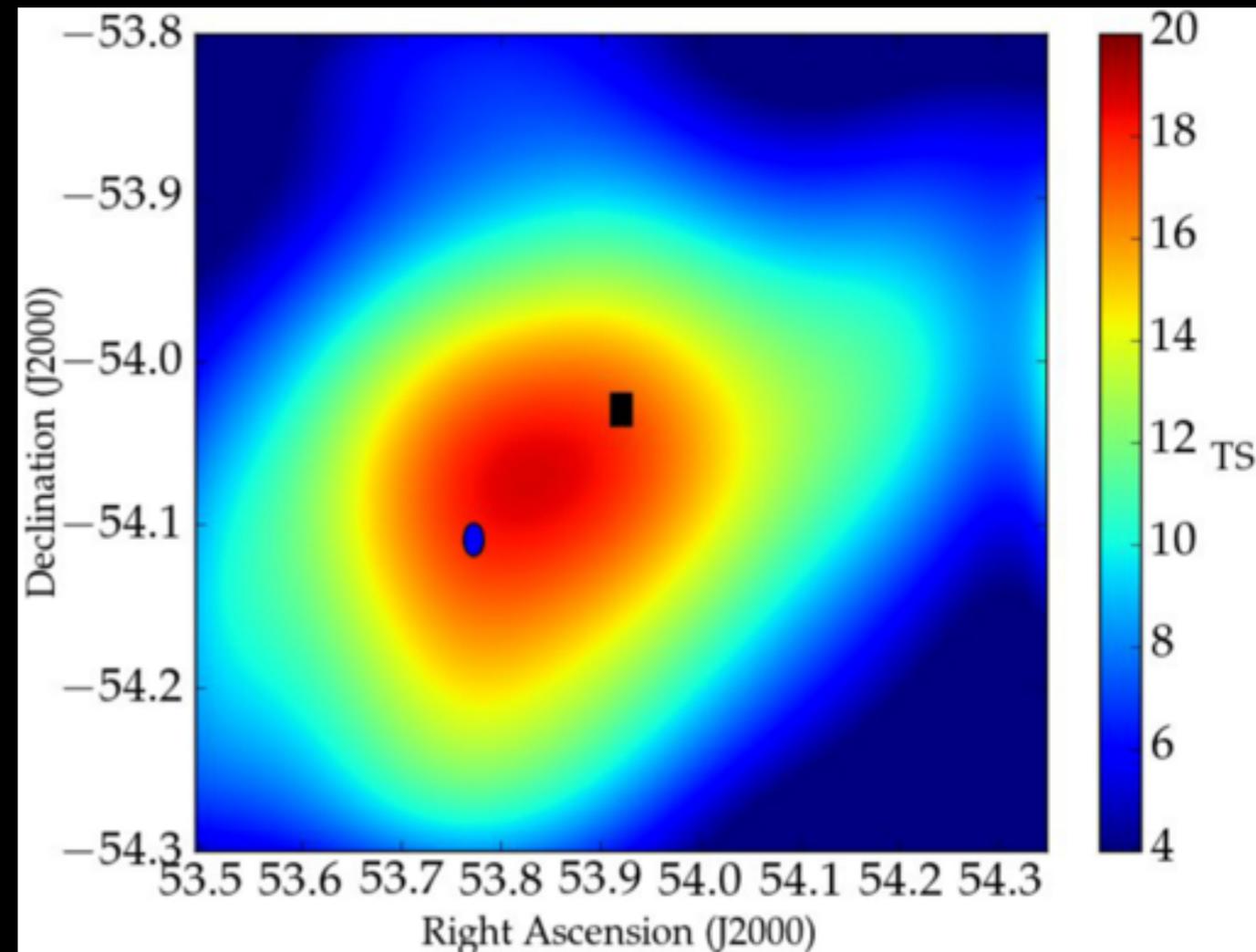
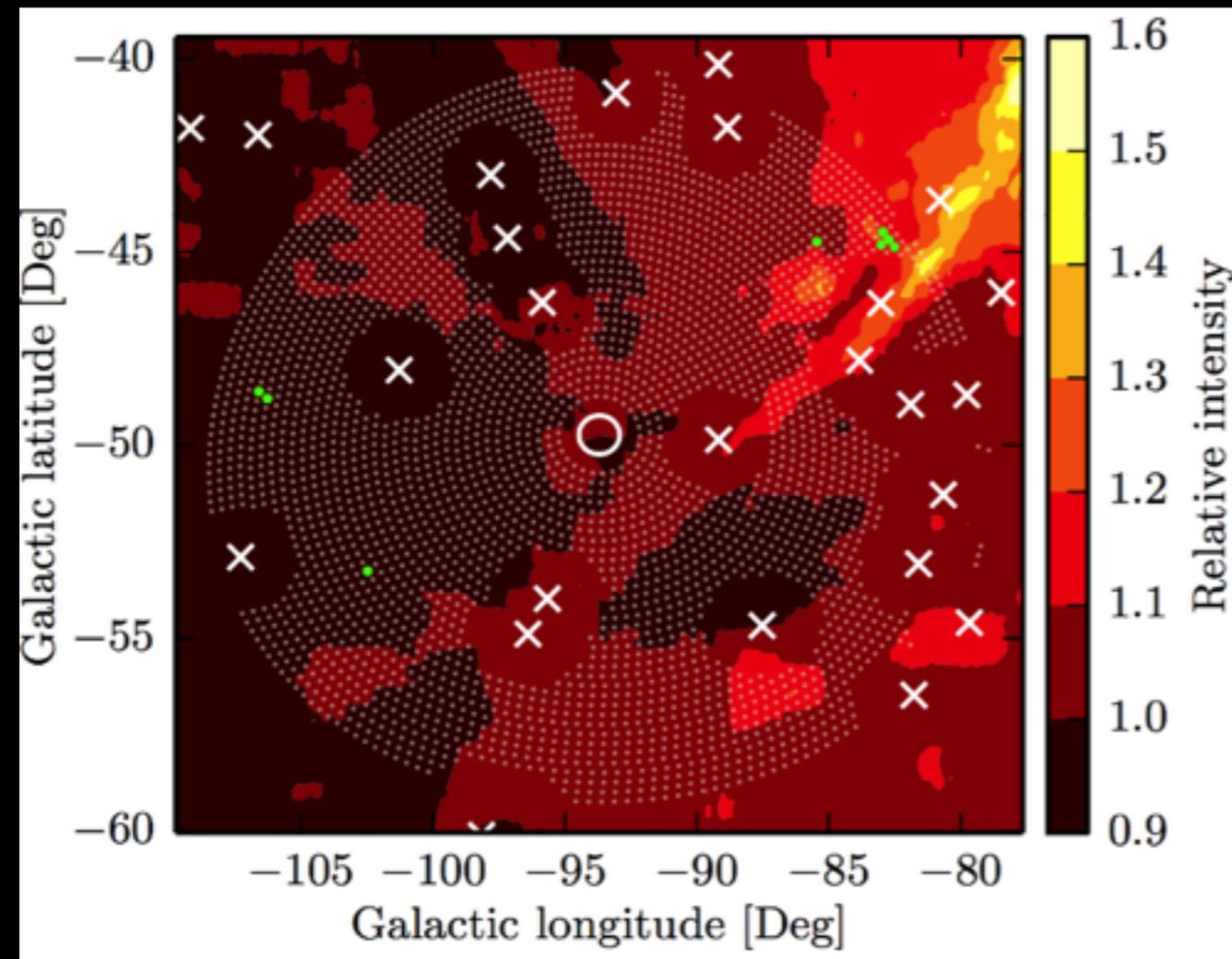


- **In the past 4 months, 12 new dwarf galaxies have been found (the first since 2009). This gives us more dark matter targets, and better sensitivity in dwarf spheroidal galaxies.**

Hot Off the Presses!

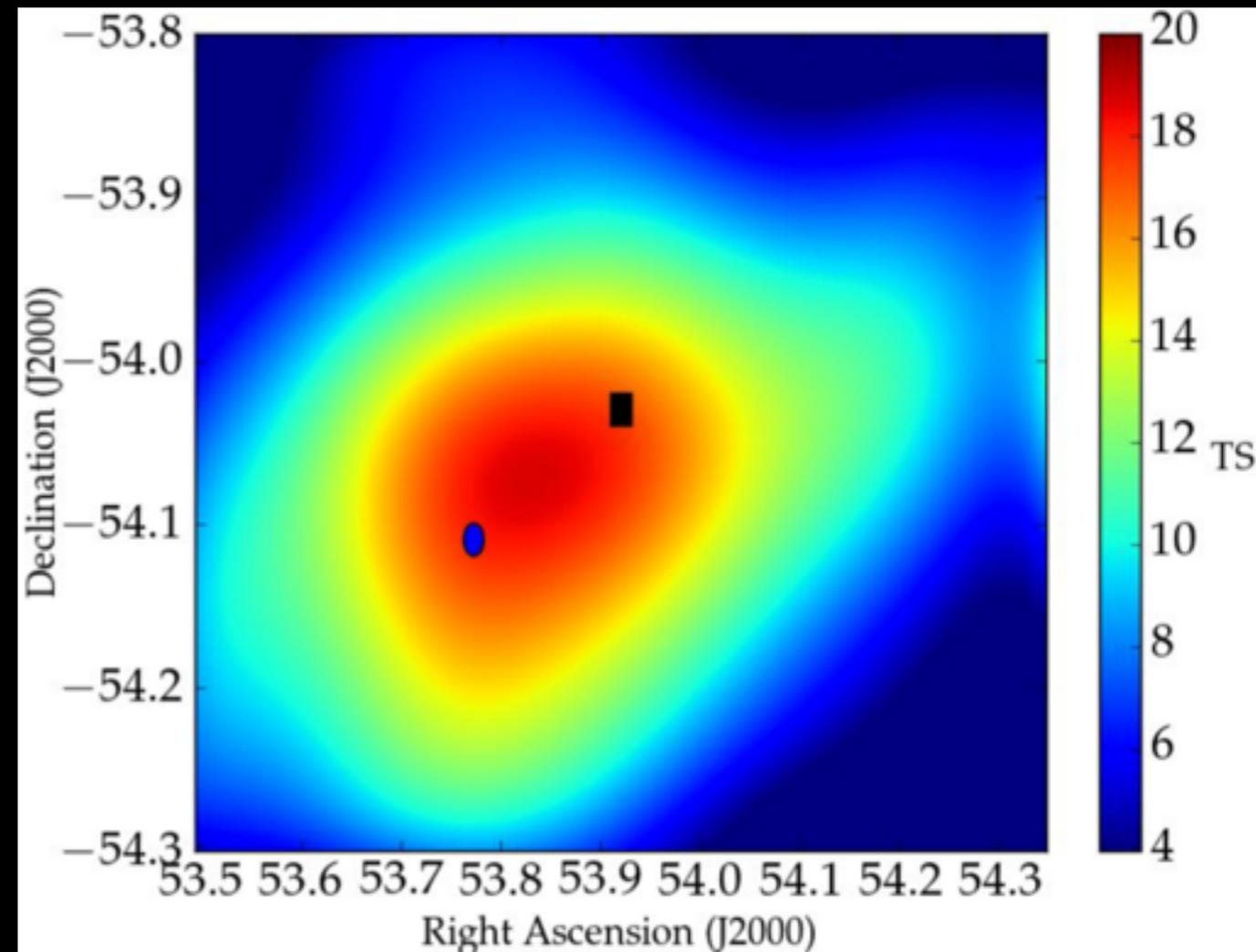
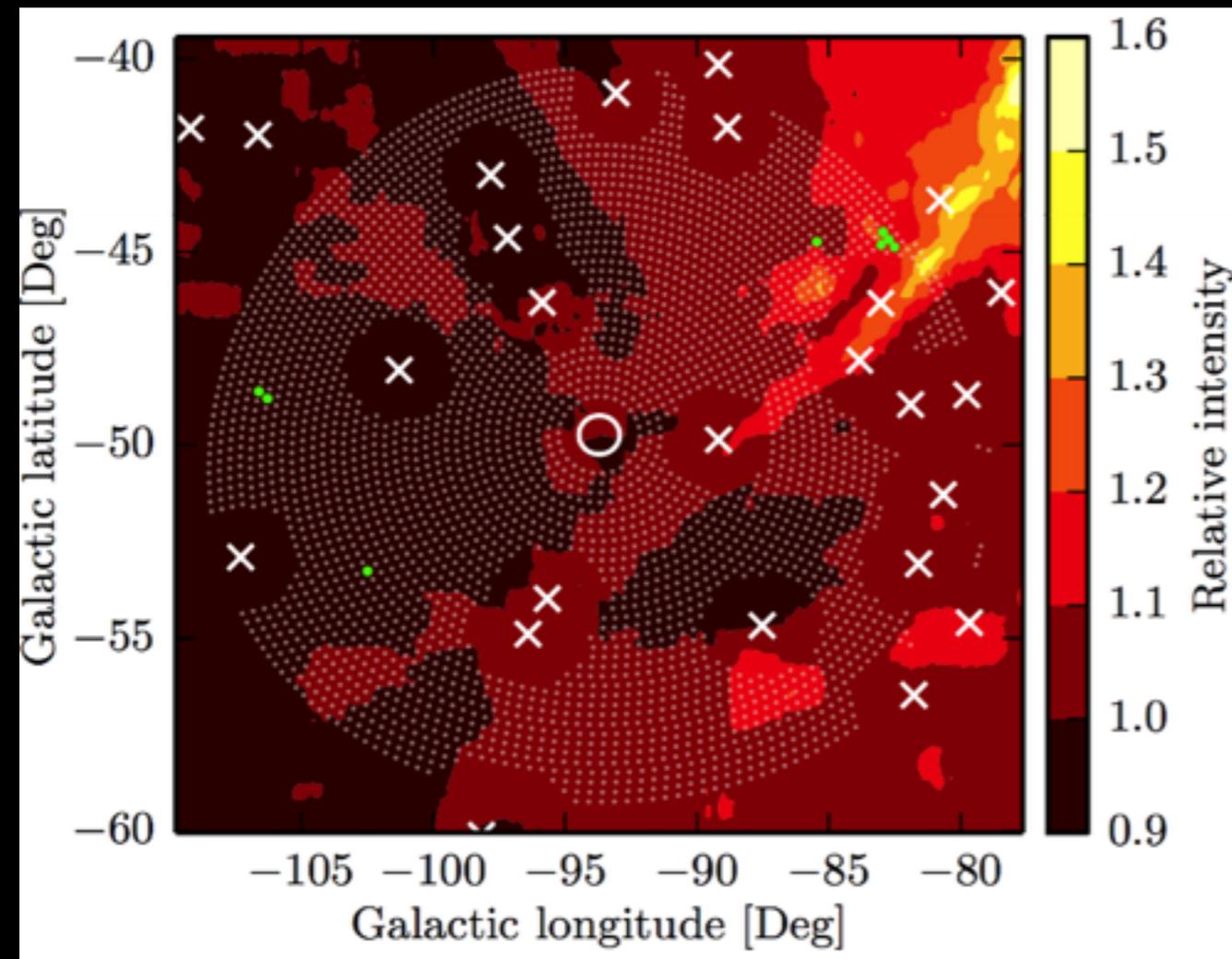


Hot Off the Presses!



- Most importantly for dark matter searches, one of these dwarfs, Reticulum 2, is one of the closest dwarfs to the sun - and thus should contain one of the brightest signals.

Hot Off the Presses!



- **And it contains a slight excess!**

Conclusions

- **The observation of fast moving stars in our galaxy has taught us a lot about how the universe evolves.**
- **We don't know what dark matter is yet - but we know a lot of the properties that the dark matter particle must have.**
- **We are just now reaching the sensitivity to explore many of the best theoretically motivated models for the dark matter particle.**