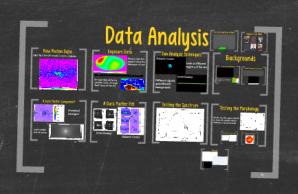
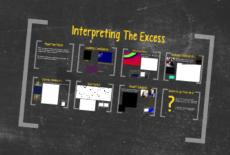
The Indirect Detection of Dark Matter in the Galactic Center







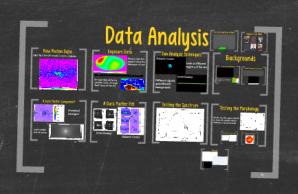
Tim Linden

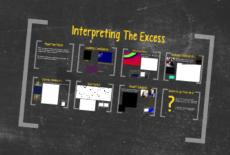
Lecture 9

Fall 2014 Compton Lectures

The Indirect Detection of Dark Matter in the Galactic Center





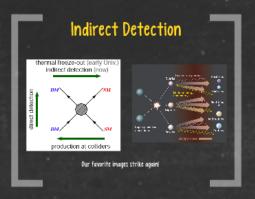


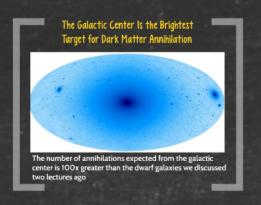
Tim Linden

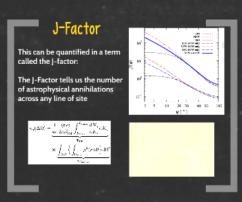
Lecture 9

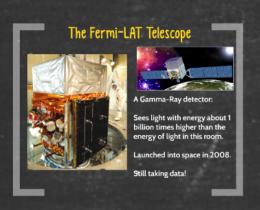
Fall 2014 Compton Lectures

Why the Galactic Center?

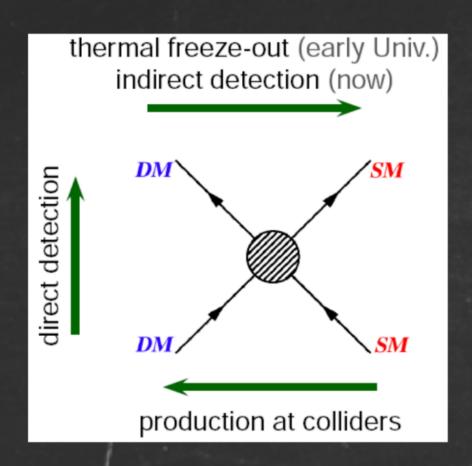


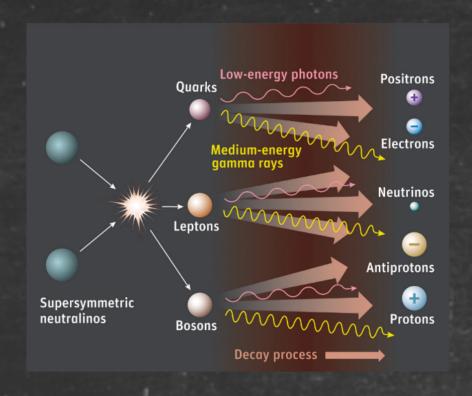






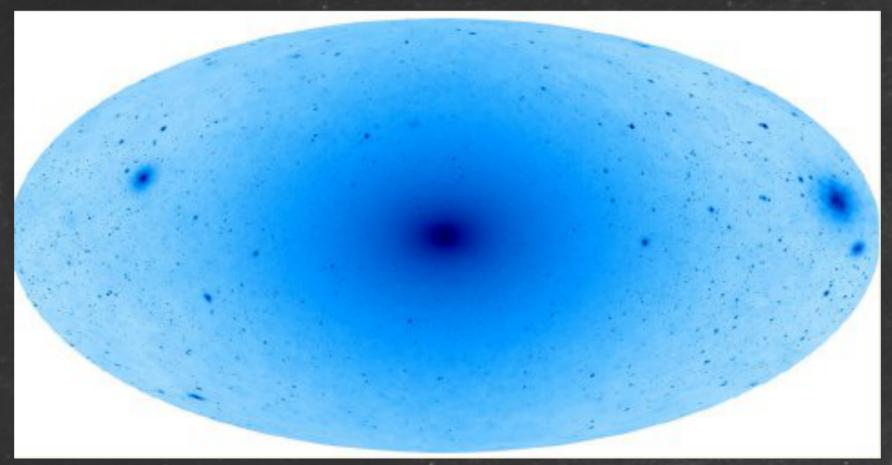
Indirect Detection





Our favorite images strike again!

The Galactic Center Is the Brightest Target for Dark Matter Annihilation



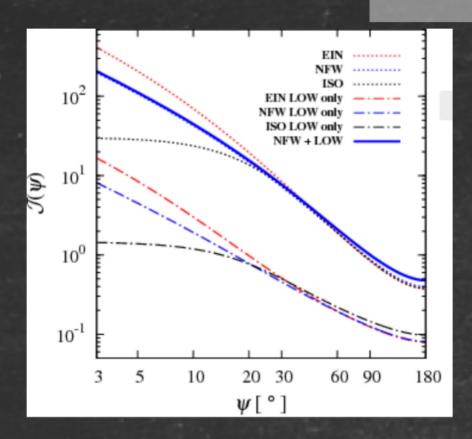
The number of annihilations expected from the galactic center is 100x greater than the dwarf galaxies we discussed two lectures ago

J-Factor

This can be quantified in a term called the J-factor:

The J-Factor tells us the number of astrophysical annihilations across any line of site

$$\phi_s(\Delta\Omega) = \underbrace{\frac{1}{4\pi} \frac{\langle \sigma v \rangle}{2m_{\rm DM}^2} \int_{E_{\rm min}}^{E_{\rm max}} \frac{{\rm d}N_{\gamma}}{{\rm d}E_{\gamma}} {\rm d}E_{\gamma}}_{\Phi_{\rm PP}} \times \underbrace{\int_{\Delta\Omega} \left\{ \int_{\rm l.o.s.} \rho^2(\boldsymbol{r}) {\rm d}l \right\} {\rm d}\Omega'}_{\text{J-factor}}$$





The Fermi-LAT Telescope





A Gamma-Ray detector:

Sees light with energy about 1 billion times higher than the energy of light in this room.

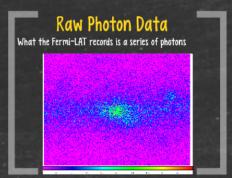
Launched into space in 2008.

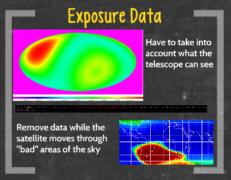
Still taking data!

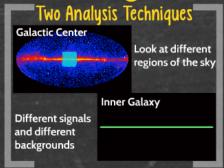
Data Analysis



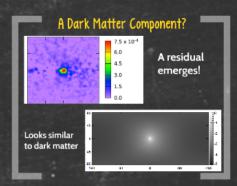


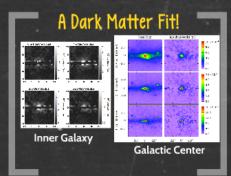


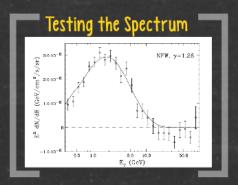


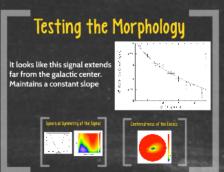














Part of a Long Stream of Papers

 Goodenough & Hooper (2009) 	arXiv:0910.2998
--	-----------------

 Hooper & Goodenough (2011, PLB 697 412) 	arXiv:1010.2752
---	-----------------

 Hooper & Linden (2011, PRD 84 12) arXiv:11 	10.0006
--	---------

- Abazajian & Kaplinghat (2012, PRD 86 8) arXiv:1207.6047
- Hooper & Slatyer (2013, PDU 2 118) arXiv:1302.6589
- Gordon & Macias (2013, PRD 88 8) arXiv:1306.5725
- Macias & Gordon (2014, PRD 89 6) arXiv:1312.6671
- Abazajian et al. (2014, PRD 90 2) arXiv:1402.4090
- Daylan et al. (2014) arXiv:1402.6703
- Calore et al. (2014) arXiv:1409.0042

Daylan et al. (2014)



Tansu Daylan



Doug Finkbeiner



Dan Hooper



Stephen Portillo



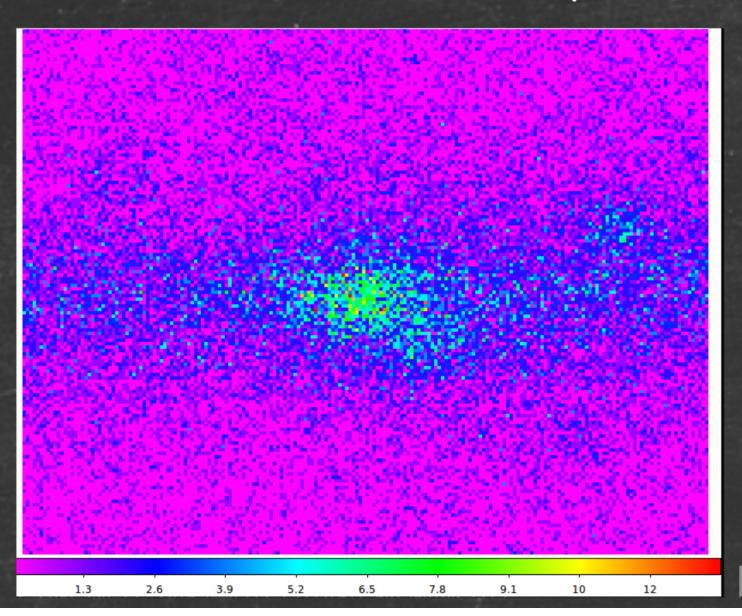
Nicholas Rodd



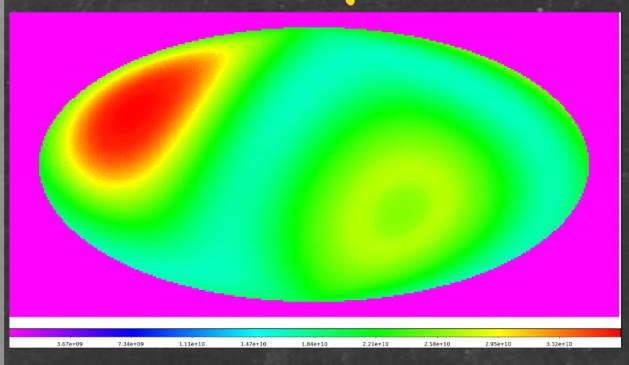
Tracy Slatyer

Raw Photon Data

What the Fermi-LAT records is a series of photons



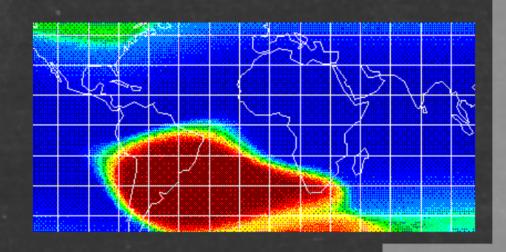
Exposure Data



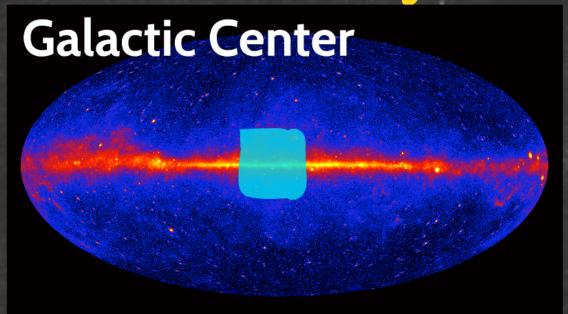
Have to take into account what the telescope can see

hice -n 10 gtmktime ft2.\$fileintime.\$fileouttime.fits "DATA_QUAL==1 && LAT_CONFIG==1 && ABS(ROCK_ANGLE)<\$maxrocking" no cft1.\$fileintime.\$fileouttime.fits nmkcft1.\$fileintime \$fileouttime.fits

Remove data while the satellite moves through "bad" areas of the sky

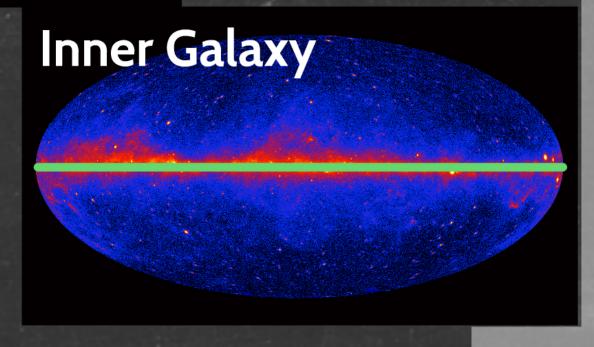


Two Analysis Techniques



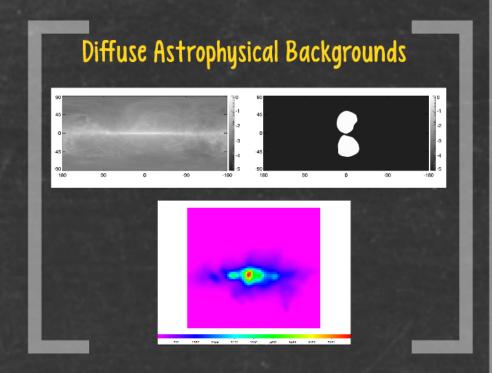
Look at different regions of the sky

Different signals and different backgrounds

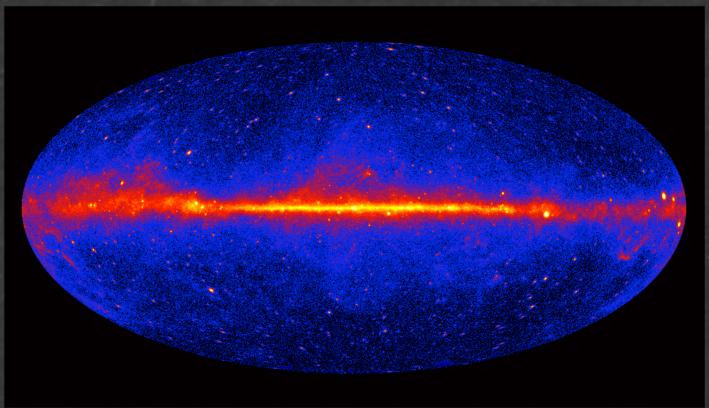


Backgrounds

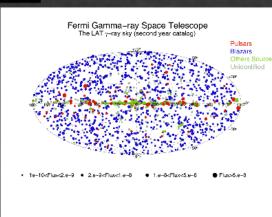
Fermi-LAT has detected more than 2300 point sources - those aren't dark matter



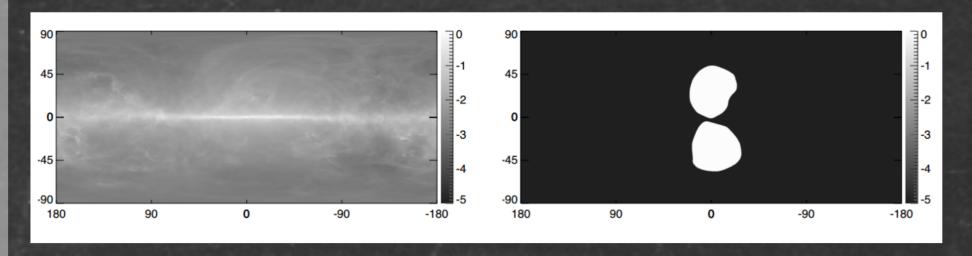
Point Source Backgrounds

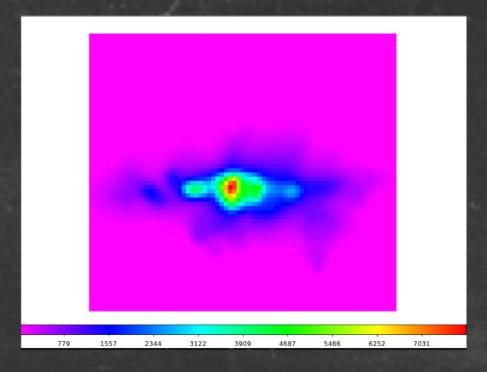


Fermi-LAT has detected more than 2300 point sources - those aren't dark matter

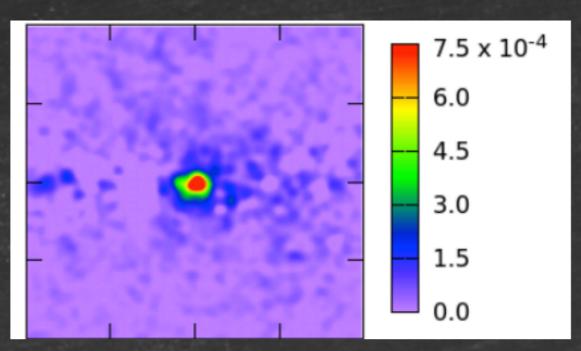


Diffuse Astrophysical Backgrounds



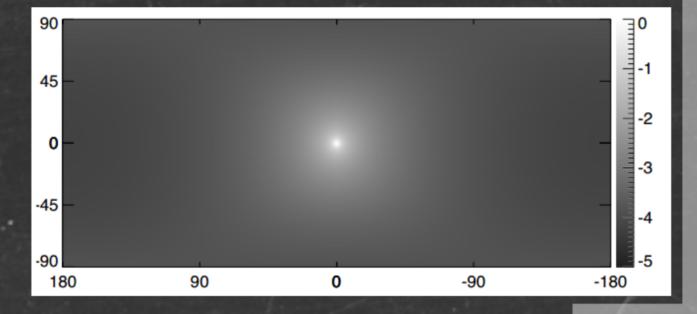


A Dark Matter Component?

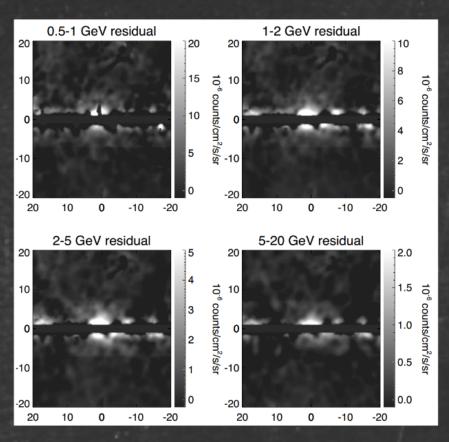


A residual emerges!

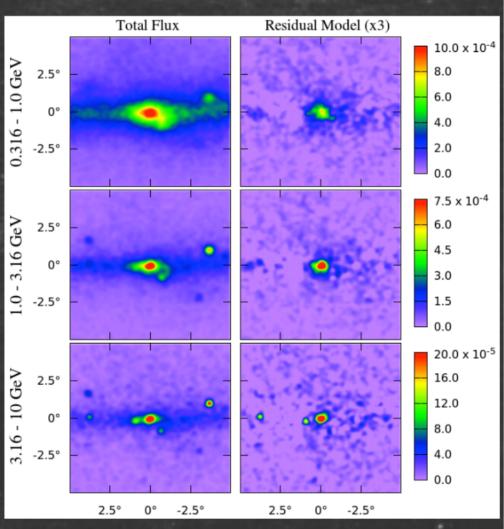
Looks similar to dark matter



A Dark Matter Fit!

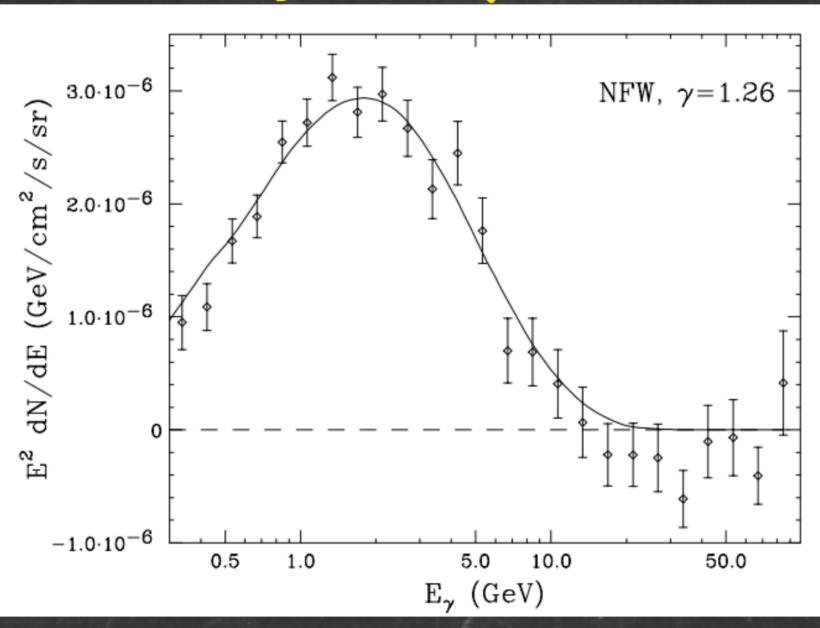


Inner Galaxy



Galactic Center

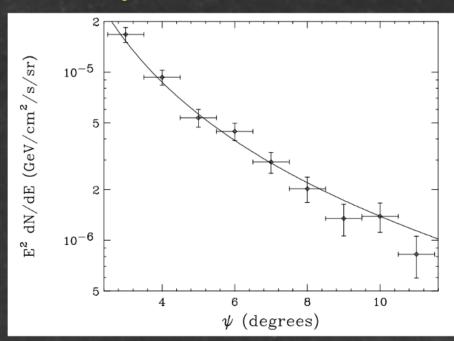
Testing the Spectrum

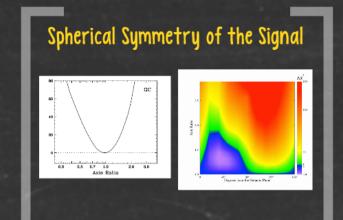


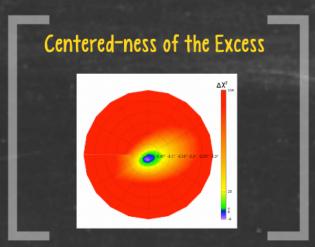
Testing the Morphology

It looks like this signal extends far from the galactic center.

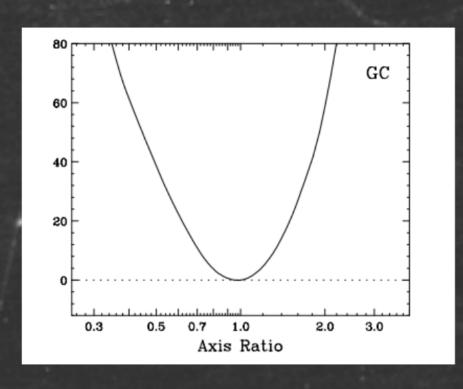
Maintains a constant slope

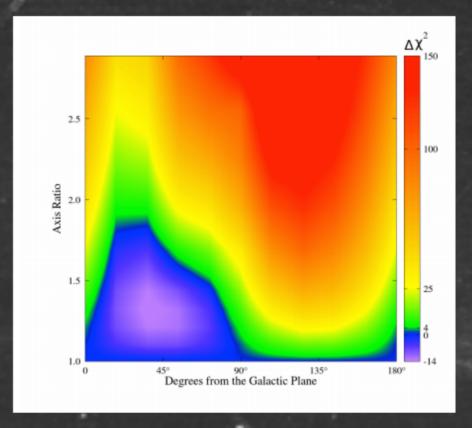




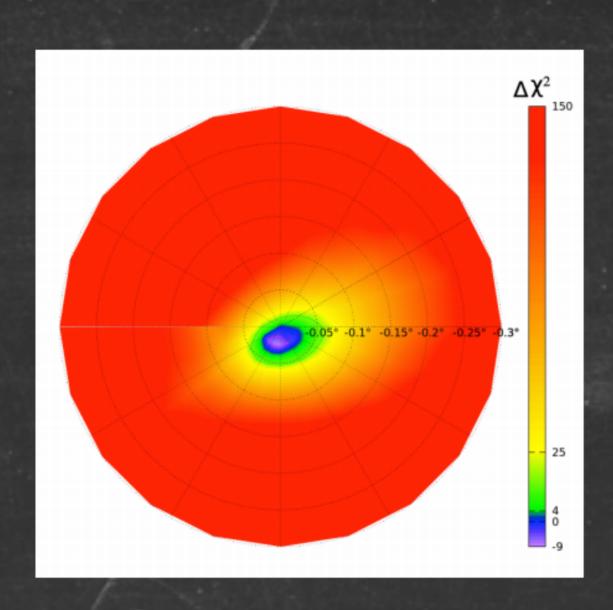


Spherical Symmetry of the Signal

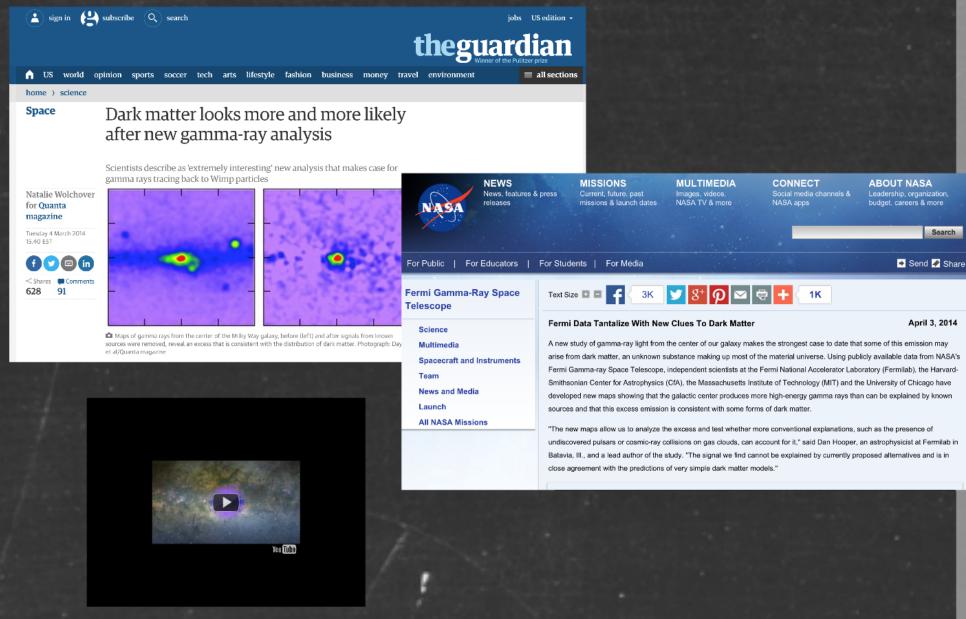




Centered-ness of the Excess



Press!



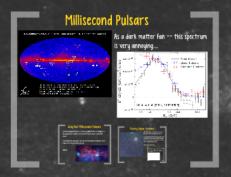
Interpreting The Excess

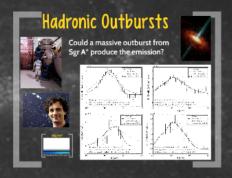
Important Note

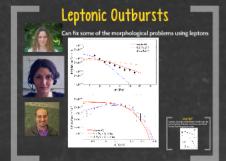
When we have said "dark matter improves the fit" what we really mean is that - "a template motivated by dark matter improves the fit"

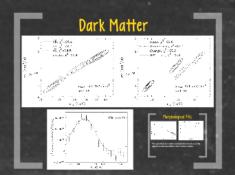
We now know there is an additional component - we just need to figure out what it is!













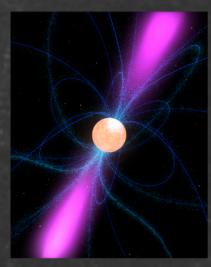


Important Note

When we have said "dark matter improves the fit" - what we really mean is that - "a template motivated by dark matter improves the fit"

We now know there is an additional component - we just need to figure out what it is!

Leading Candidates



Pulsars

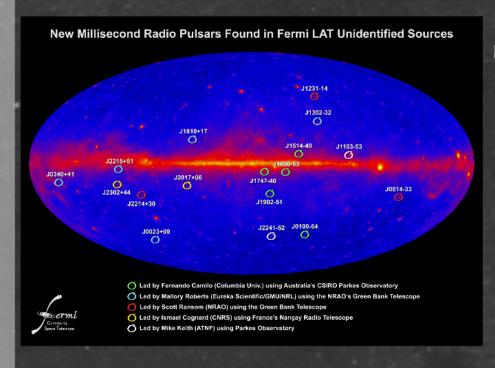


Outbursts from The Galactic Center

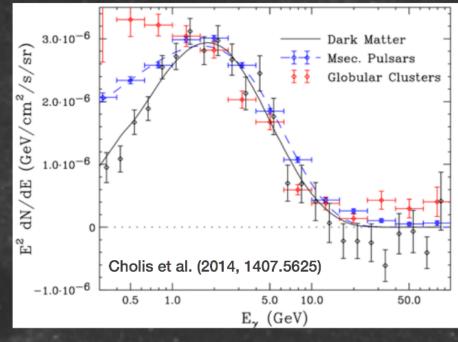


Dark Matter

Millisecond Pulsars



As a dark matter fan -- this spectrum is very annoying....







Why Not Millisecond Pulsars

If Millisecond Pulsars make up all of the emission in the galactic center, we should have seen some by now.

Instead, we actually see no millisecond pulsars very close to the galactic center.

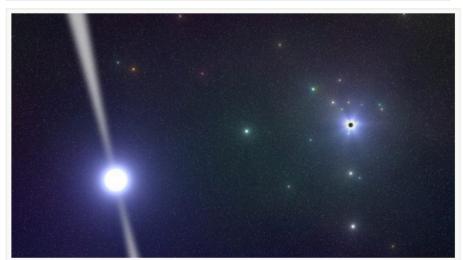
Why?



Missing Pulsar Problem

Astronomers Discover a Magnetar at the Galactic Center

August 16, 2013 Space



Artist's impression of PSR J1745-2900, a pulsar with a very high magnetic field ("magnetar") in direct vicinity of the central source of our Galaxy, a supermassive black hole of approximately 4 million times the mass of our sun. Measurements of the pulsar imply that a strong magnetic field exists in the vicinity around the black hole. MPIfR/Ralph Eatough.

An international team of astronomers has discovered a magnetar at the center of the Milky Way galaxy.

Astronomers have discovered a magnetar at the center of our Milky Way. This pulsar has an extremely strong magnetic field and enables researchers to investigate the direct vicinity of the black hole at the heart of the galaxy. An international team of scientists

Up until 2013, no pulsars had been discovered at the galactic center.

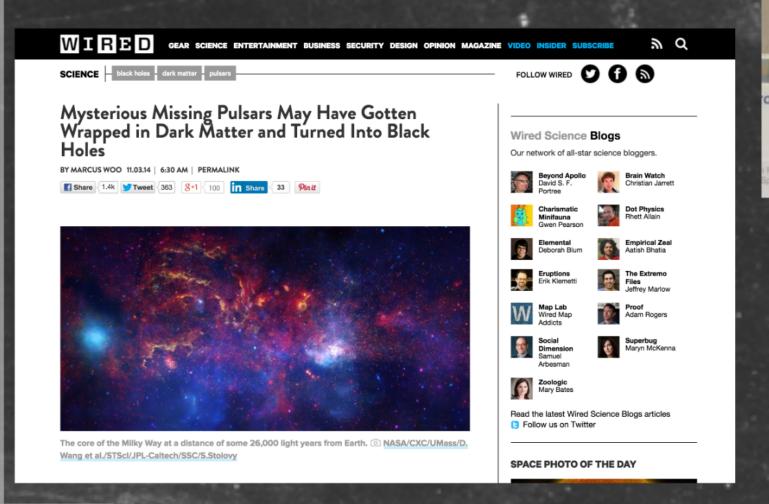
It was possible that the high density of electrons near the galactic center was smearing out the signal, making them impossible to see.

Then we found one.....



Another Sign of Dark Matter?

One one (very young) pulsar, called a magnetar has been discovered near the galactic center. We expect there to be thousands.....



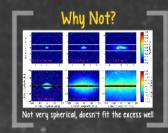
Hadronic Outbursts

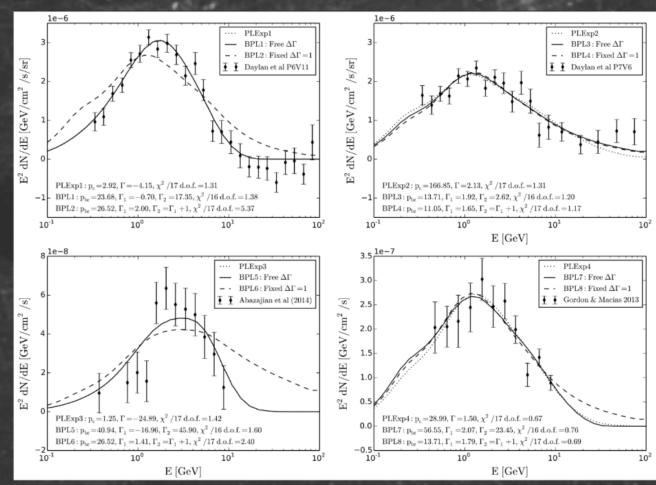


Could a massive outburst from Sgr A* produce the emission?

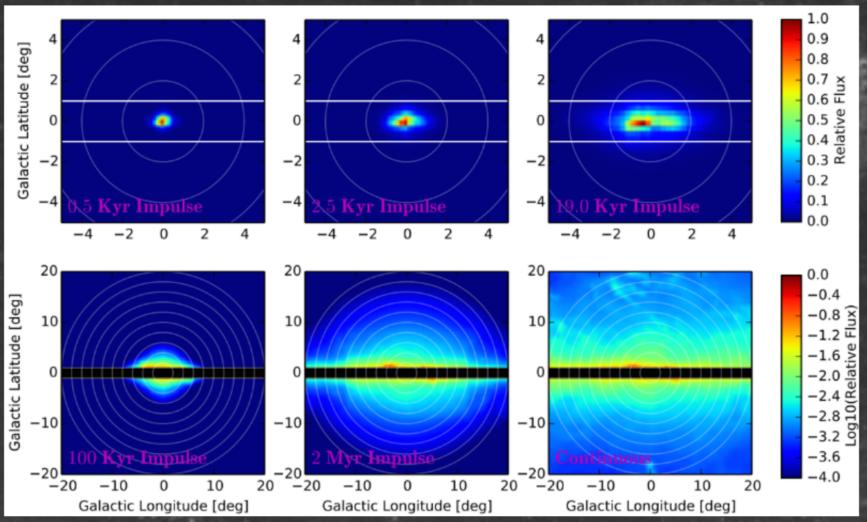








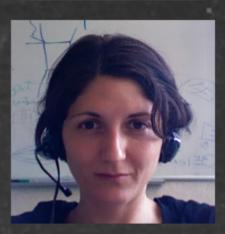
Why Not?



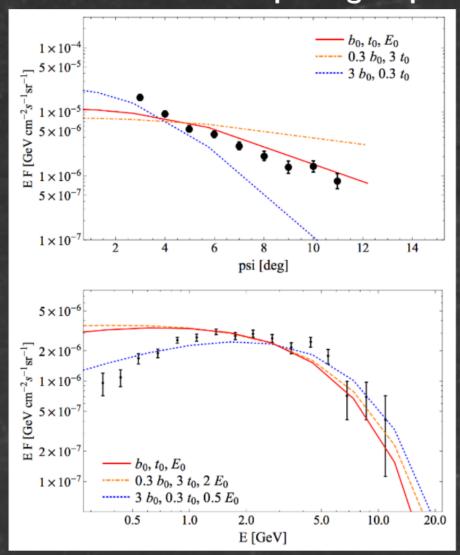
Not very spherical, doesn't fit the excess well

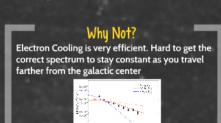
Leptonic Outbursts

Can fix some of the morphological problems using leptons



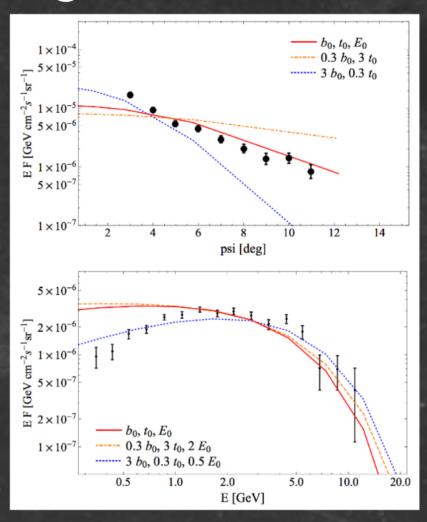




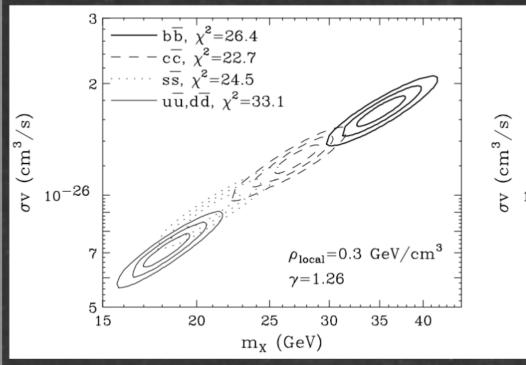


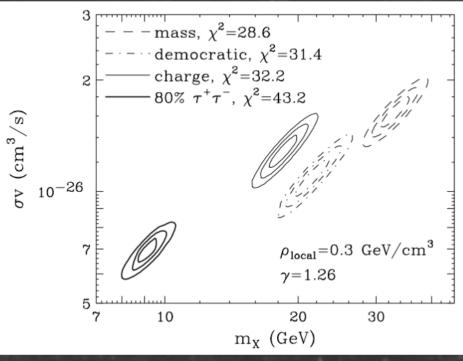
Why Not?

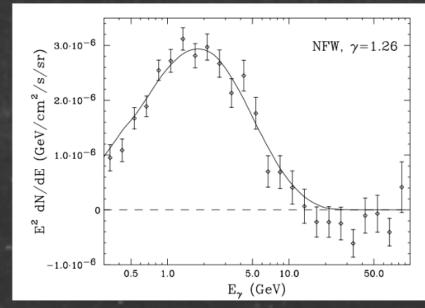
Electron Cooling is very efficient. Hard to get the correct spectrum to stay constant as you travel farther from the galactic center

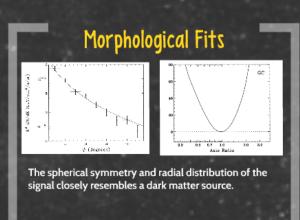


Dark Matter

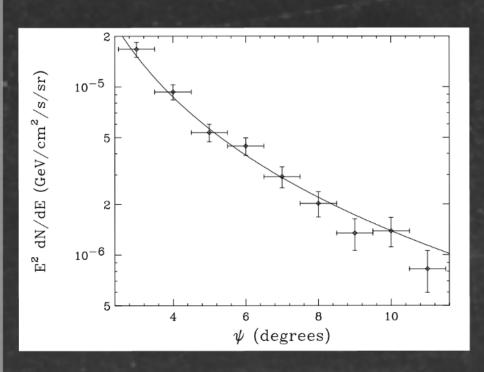


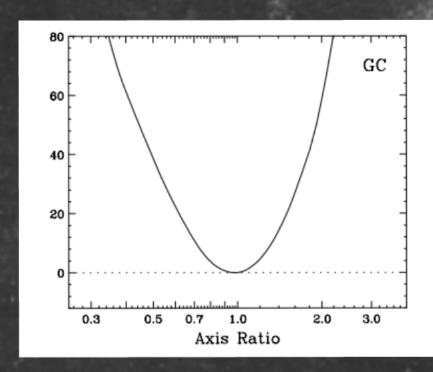






Morphological Fits

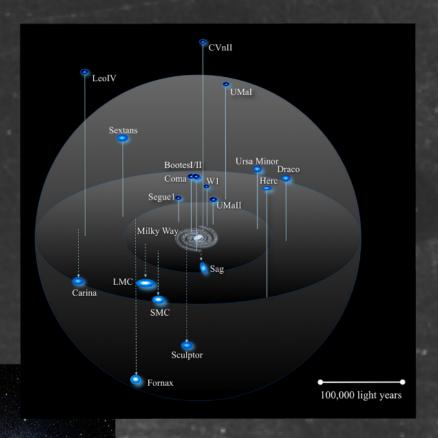




The spherical symmetry and radial distribution of the signal closely resembles a dark matter source.

Dwarf Galaxies

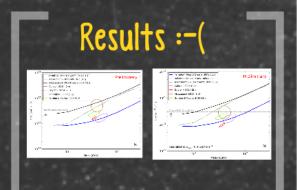
Fermi dwarf galaxies could provide a strong confirmation of the signal



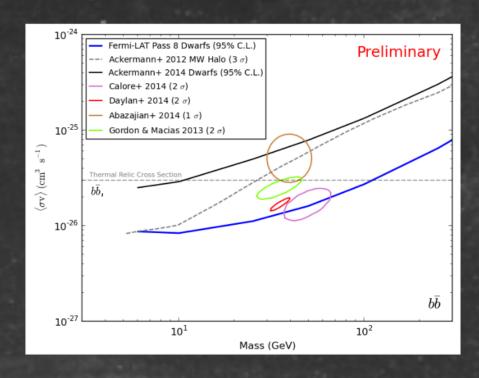


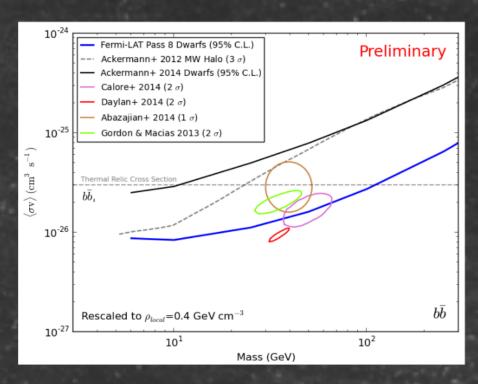






Results :-(





Where to go From Here?

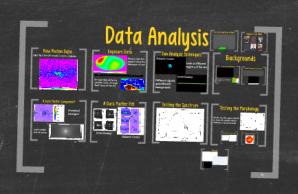


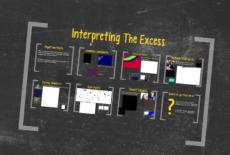
"Our case is very much a process-of-elimination argument. We made a list, scratched off things that didn't work, and ended up with dark matter," - Doug Finkbeiner

If we knew what it was we were doing, it would not be called research, would it?" - Albert Einstein

The Indirect Detection of Dark Matter in the Galactic Center







Tim Linden

Lecture 9

Fall 2014 Compton Lectures